

École Doctorale des Sciences de l'Environnement d'Île-de-France

Année Universitaire 2012-2013

Modélisation Numérique
de l'Écoulement Atmosphérique
et Assimilation d'Observations

Olivier Talagrand

Cours 8

31 Mai 2013

How to write the adjoint of a code ?

Operation $a = b \times c$

Input b, c

Output a but also b, c

For clarity, we write

$$a = b \times c$$

$$b' = b$$

$$c' = c$$

$\partial J / \partial a$, $\partial J / \partial b'$, $\partial J / \partial c'$ available. We want to determine $\partial J / \partial b$, $\partial J / \partial c$

Chain rule

$$\partial J / \partial b = (\partial J / \partial a) \underset{c}{(\partial a / \partial b)} + (\partial J / \partial b') \underset{1}{(\partial b' / \partial b)} + (\partial J / \partial c') \underset{0}{(\partial c' / \partial b)}$$

$$\partial J / \partial b = (\partial J / \partial a) c + \partial J / \partial b'$$

Similarly

$$\partial J / \partial c = (\partial J / \partial a) b + \partial J / \partial c'$$

Exact bayesian estimation ?

Particle filters

Predicted ensemble at time t : $\{x_n^b, n = 1, \dots, N\}$, each element with its own weight (probability) $P(x_n^b)$

Observation vector at same time : $y = Hx + \varepsilon$

Bayes' formula

$$P(x_n^b|y) \sim P(y|x_n^b) P(x_n^b)$$

Defines updating of weights

Bayes' formula

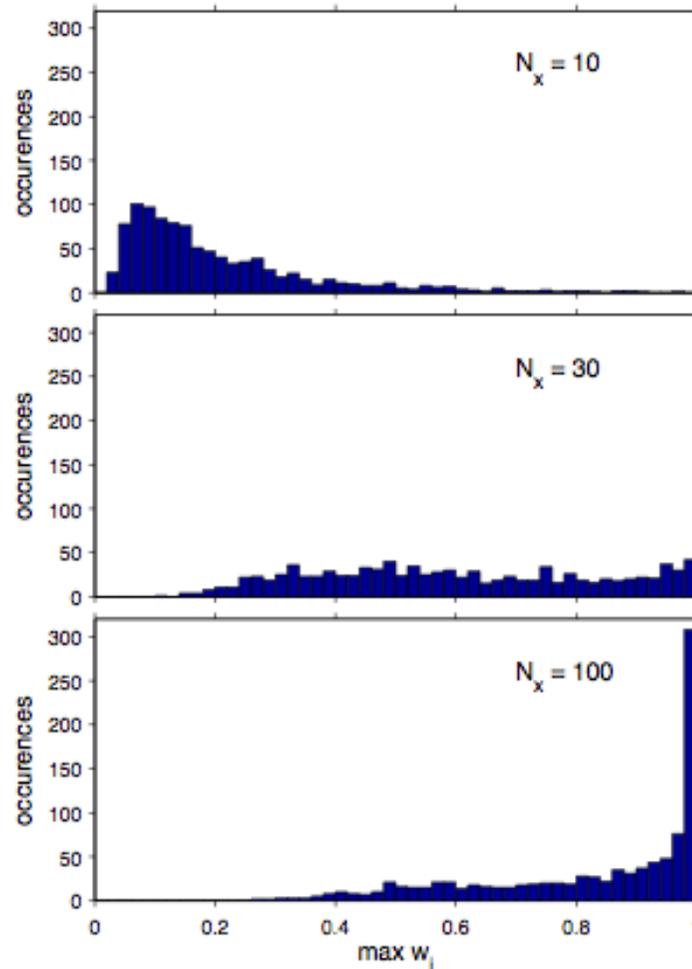
$$P(x_n^b|y) \sim P(y|x_n^b) P(x_n^b)$$

Defines updating of weights; particles are not modified. Asymptotically converges to bayesian pdf. Very easy to implement.

Observed fact. For large state dimension, ensemble tends to collapse.

Behavior of $\max w^i$

▷ $N_e = 10^3$; $N_x = 10, 30, 100$; 10^3 realizations



average squared error of
posterior mean = 5.5

... = 25

... = 127

C. Snyder, <http://www.cawcr.gov.au/staff/pxs/wmoda5/Oral/Snyder.pdf>

Problem originates in the ‘curse of dimensionality’ Large dimension pdf’s are very diffuse, so that very few particles (if any) are present in areas where conditional probability (‘*likelihood*’) $P(y|x)$ is large.

Bengtsson *et al.* (2008) and Snyder *et al.* (2008) evaluate that stability of filter requires the size of ensembles to increase exponentially with space dimension.

Alternative possibilities (review in van Leeuwen, 2009, *Mon. Wea. Rev.*, 4089-4114)

Resampling. Define new ensemble.

Simplest way. Draw new ensemble according to probability distribution defined by the updated weights. Give same weight to all particles. Particles are not modified, but particles with low weights are likely to be eliminated, while particles with large weights are likely to be drawn repeatedly. For multiple particles, add noise, either from the start, or in the form of ‘model noise’ in ensuing temporal integration.

Random character of the sampling introduces noise. Alternatives exist, such as *residual sampling* (Lui and Chen, 1998, van Leeuwen, 2003). Updated weights w_n are multiplied by ensemble dimension N . Then p copies of each particle n are taken, where p is the integer part of Nw_n . Remaining particles, if needed, are taken randomly from the resulting distribution.

Importance Sampling.

Use a *proposal density* that is closer to the new observations than the density defined by the predicted particles (for instance the density defined by EnKF, after the latter has used the new observations). Independence between observations is then lost in the computation of likelihood $P(y|x)$ (or is it not ?)

In particular, *Guided Sequential Importance Sampling* (van Leeuwen, 2002).
Idea : use observations performed at time k to resample ensemble at some timestep anterior to k , or ‘nudge’ integration between times $k-1$ and k towards observation at time k .

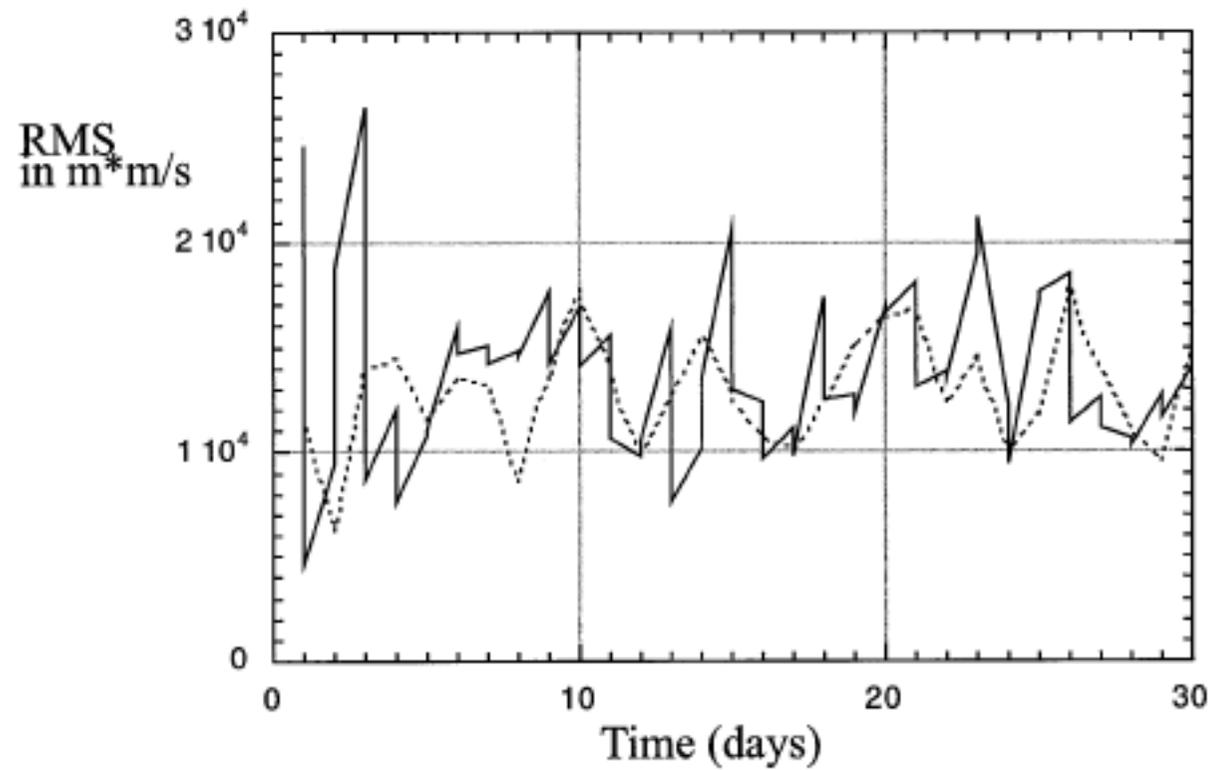
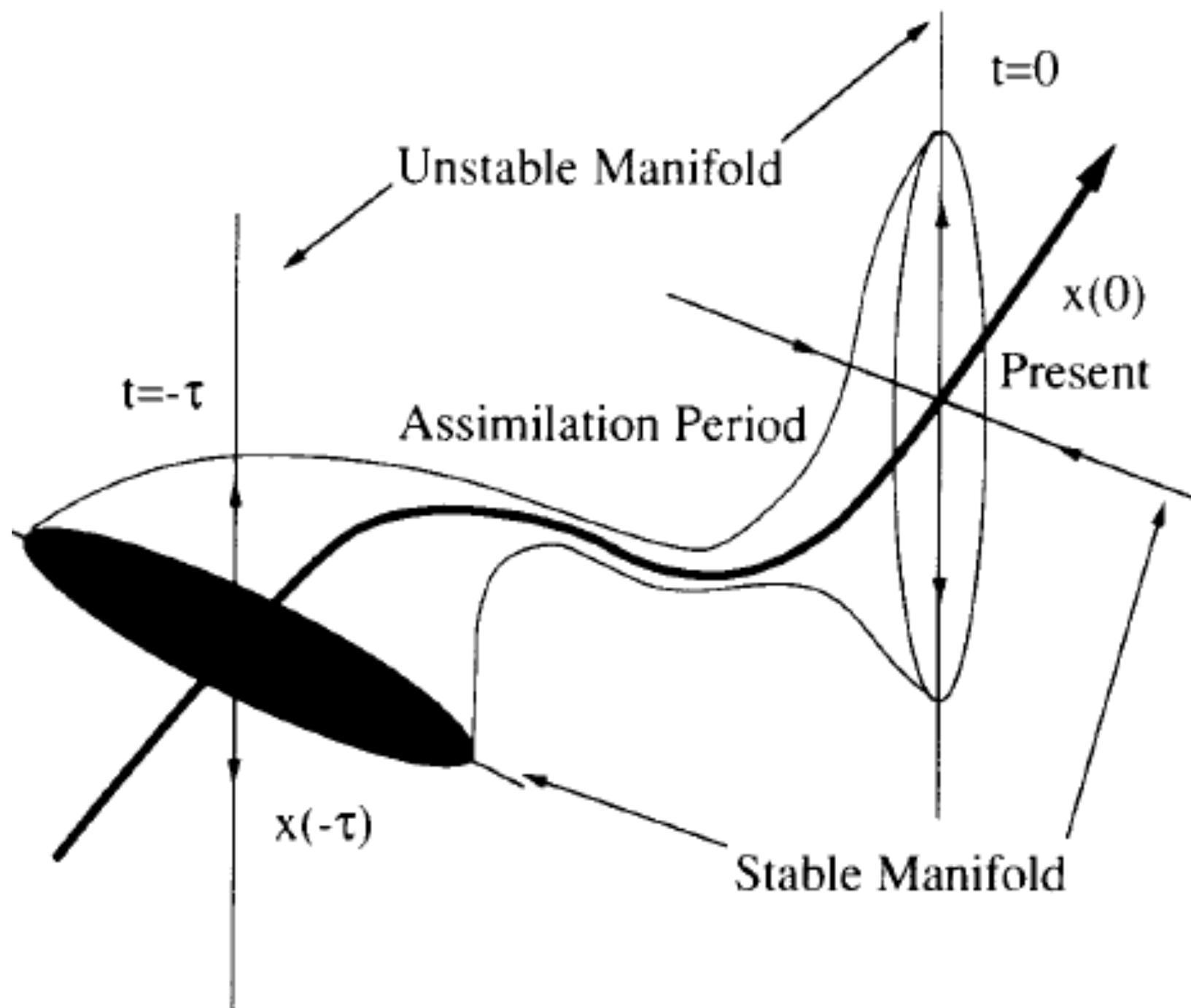


FIG. 12. Comparison of rms error ($\text{m}^2 \text{s}^{-1}$) between ensemble mean and independent observations (dotted line) and the std dev in the ensemble (solid line). The excellent agreement shows that the SIRF is working correctly.

Particle filters are actively studied (van Leeuwen, Morzfeld, ...)

If there is uncertainty on the state of the system, and dynamics of the system is perfectly known, uncertainty on the state along stable modes decreases over time, while uncertainty along unstable modes increases.

Stable (unstable) modes : perturbations to the basic state that decrease (increase) over time.



Consequence : Consider 4D-Var assimilation, which carries information both forward and backward in time, performed over time interval $[t_0, t_1]$ over uniformly distributed noisy data. If assimilating model is perfect, estimation error is concentrated in stable modes at time t_0 , and in unstable modes at time t_1 . Error is smallest somewhere within interval $[t_0, t_1]$.

Similar result holds true for Kalman filter (or more generally any form of sequential assimilation), in which estimation error is concentrated in unstable modes at any time.

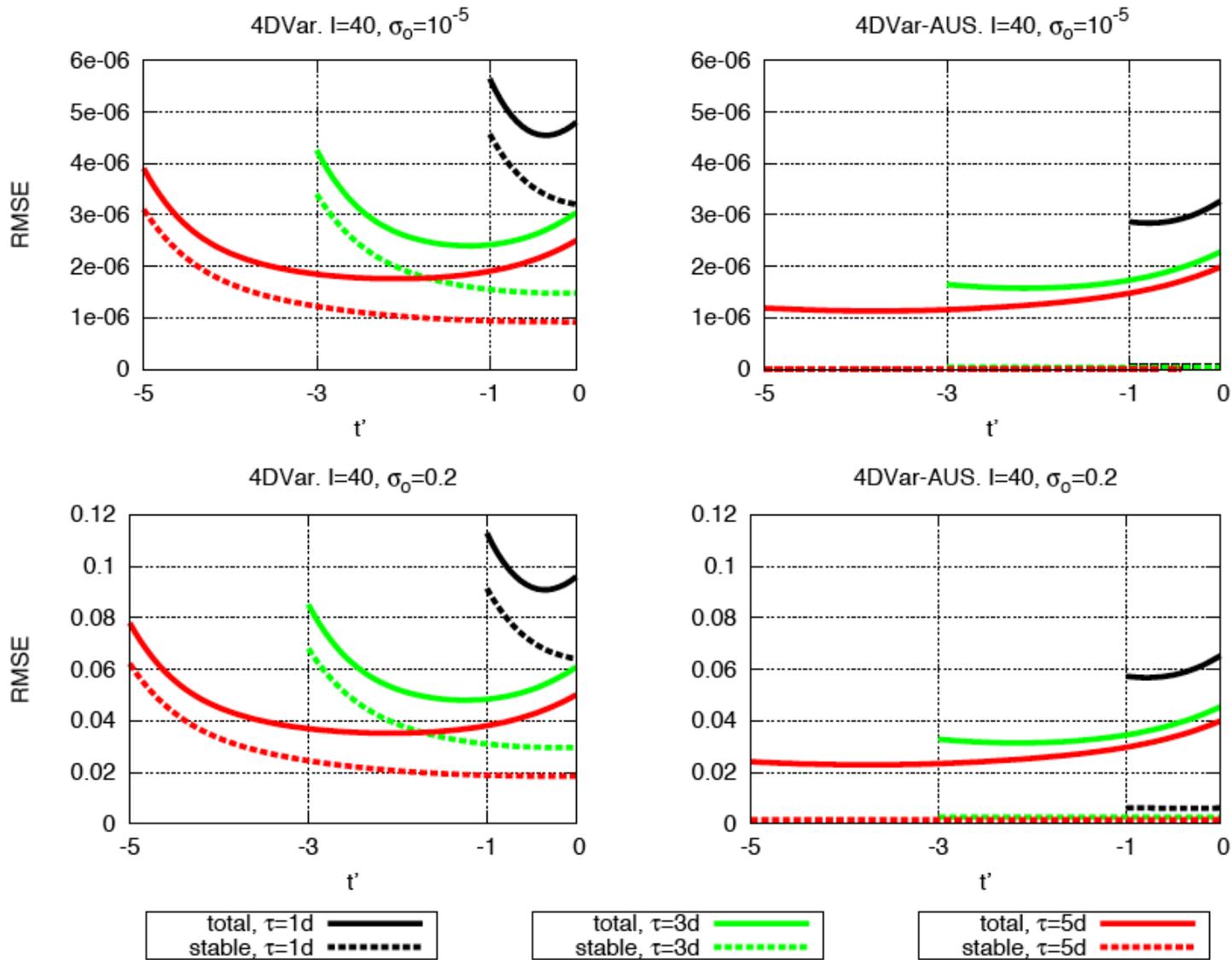


Figure 3. Time average RMS error within 1, 3, 5 days assimilation windows as a function of $t' = t - \tau$, with $\sigma_0 = .2, 10^{-5}$ for the model configuration $I = 40$. Left panel: 4DVar. Right panel: 4DVar-AUS with $N = 15$. Solid lines refer to total assimilation error, dashed lines refer to the error component in the stable subspace e_{16}, \dots, e_{40} .

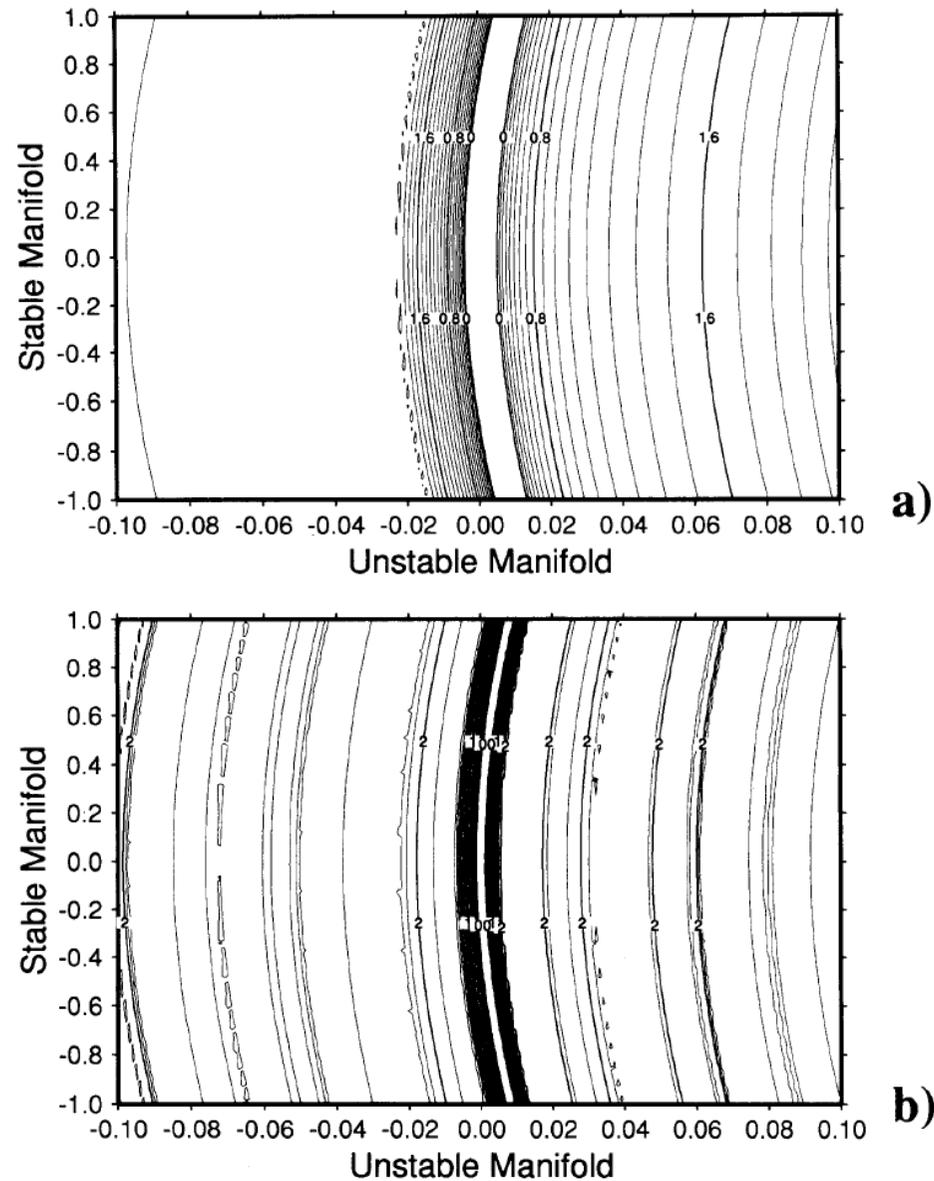


Fig. 3. Variations of the error-free forward cost-function $J'_2(\tau, \hat{x}, x)$ (Lorenz system) in the plane spanned by the stable and unstable directions, as determined from the tangent linear system (see text), and for $\tau = 6$ (panel (a)) and $\tau = 8$ (panel (b)) respectively. The metric has been distorted in order to make the stable and unstable manifolds orthogonal to each other in the figure. The scale on the contour lines is logarithmic (decimal logarithm). Contour interval: 0.1. For clarity, negative contours, which would be present only in the central "valley" directed along the stable manifold, have not been drawn.

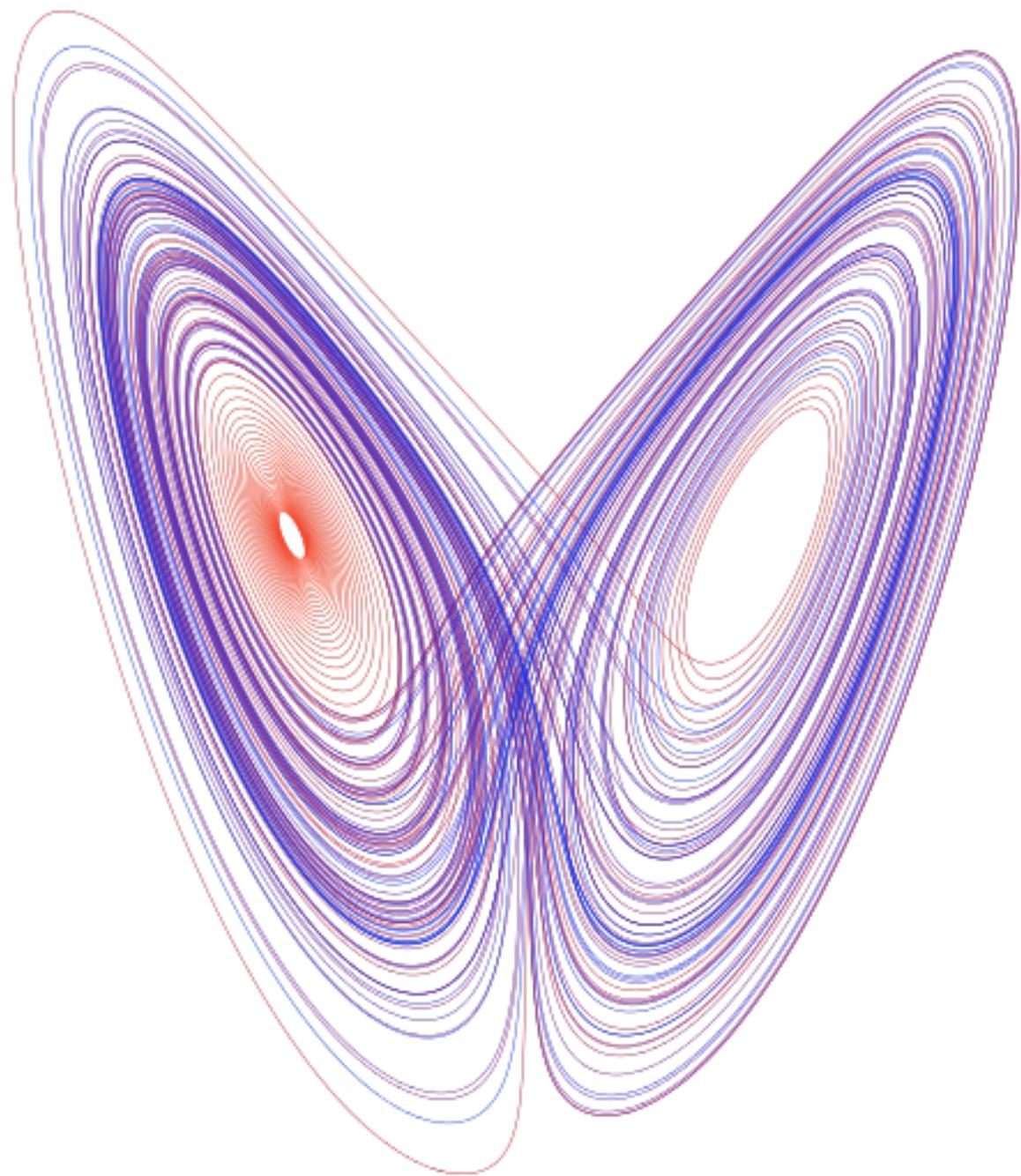
Lorenz (1963)

$$dx/dt = \sigma(y-x)$$

$$dy/dt = \rho x - y - xz$$

$$dz/dt = -\beta z + xy$$

with parameter values $\sigma = 10$, $\rho = 28$, $\beta = 8/3 \Rightarrow$ chaos



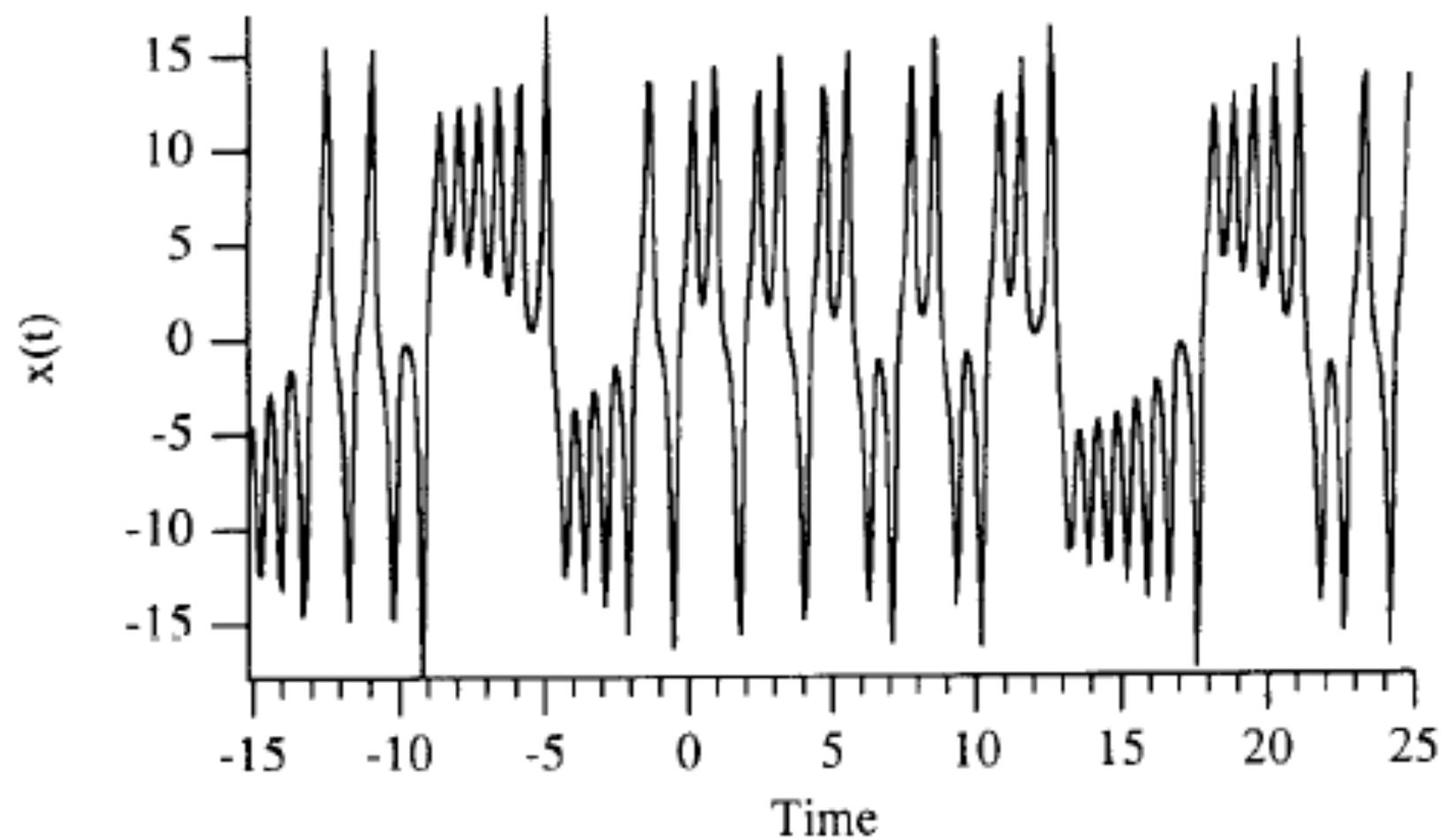


Fig. 2. Time variations, along the reference solution, of the variable $x(t)$ of the Lorenz system.

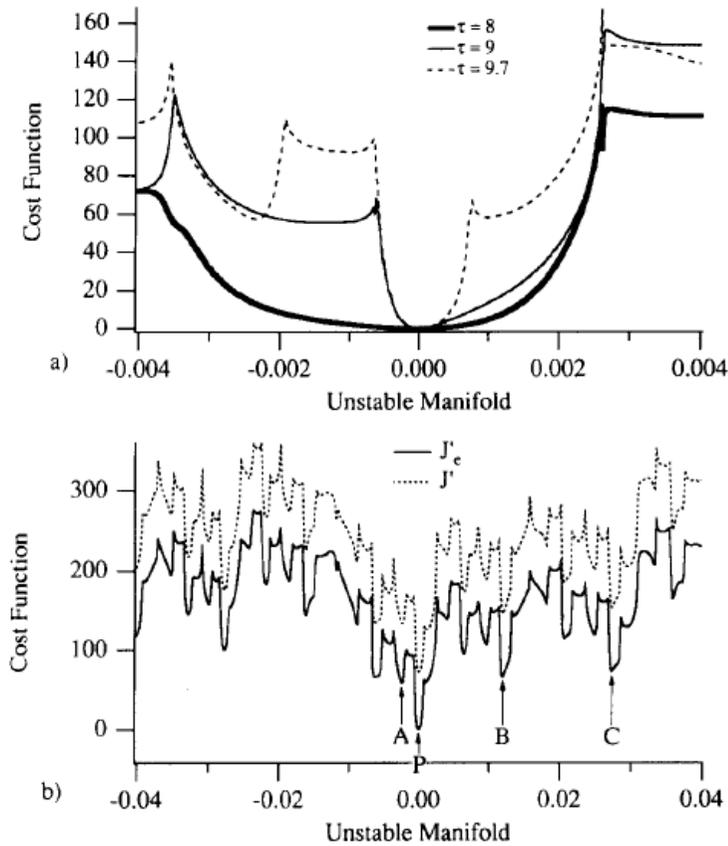


Fig. 4. Panel (a): Cross-section of the error-free forward cost-function $J'_e(\tau, \hat{x}, x)$ along the unstable manifold, for various values of τ . Panel (b). As in panel (a), for $\tau = 9.7$, and with a display interval ten times as large, respectively for the error-free forward cost-function $J'_e(\tau, \hat{x}, x)$ (solid curve) and for the error-contaminated cost-function $J_e(\tau, \hat{x}, x)$ (dashed curve). In the latter case, the total variance of the observational noise is $E^2 = 75$.

Pires *et al.*, *Tellus*, 1996 ; Lorenz system (1963)

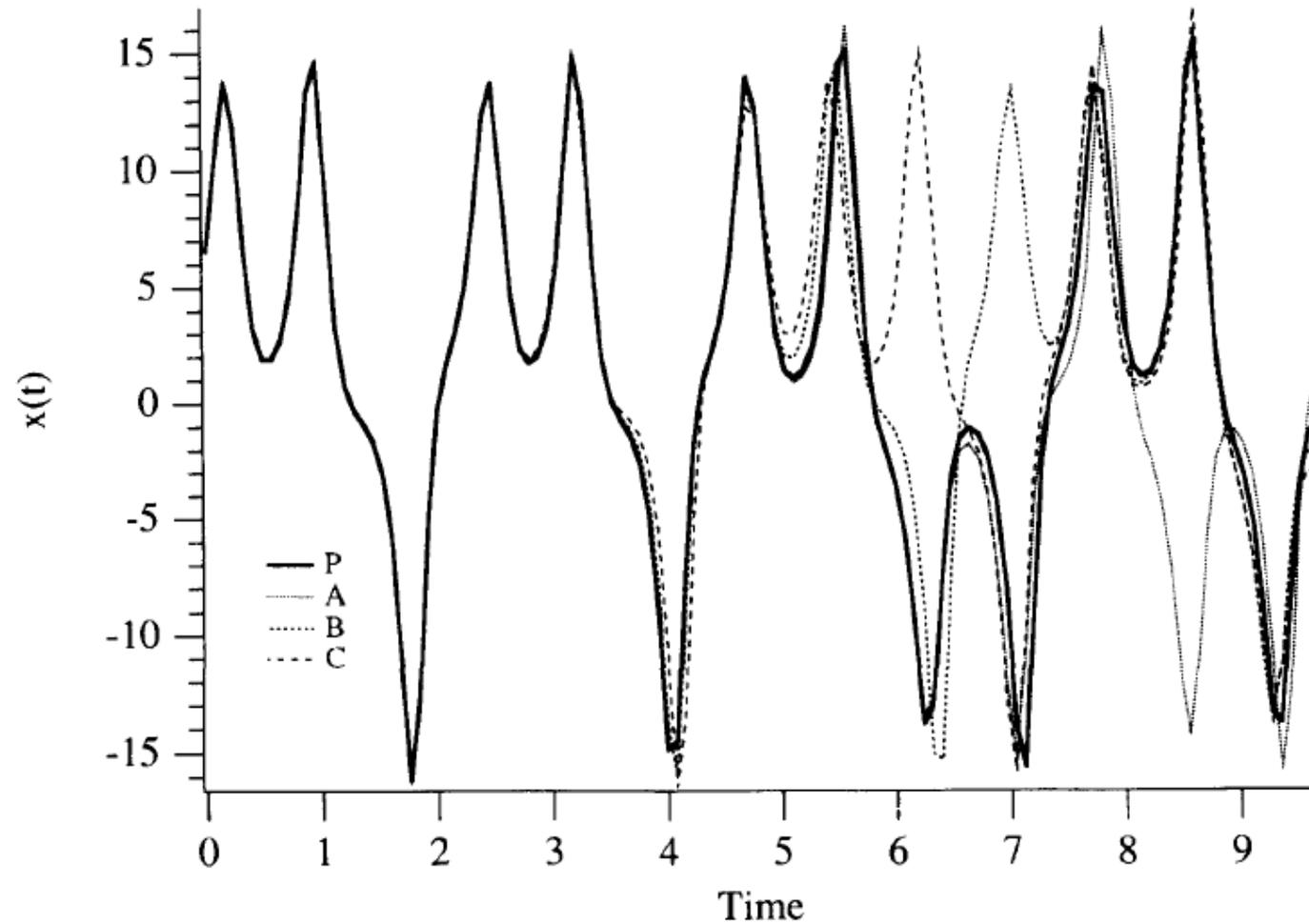


Fig. 5. Variations of the coordinate x along the orbits originating from the minima P , A , B , C (indicated in Fig. 4b) of the error-free cost-function.

Minima in the variations of objective function correspond to solutions that have bifurcated from the observed solution, and to different folds in state space.

Quasi-Static Variational Assimilation (QSVA). Increase progressively length of the assimilation window, starting each new assimilation from the result of the previous one. This should ensure, at least if observations are in a sense sufficiently dense in time, that current estimation of the system always lies in the attractive basin of the absolute minimum of objective function (Pires *et al.*, Swanson *et al.*, Luong, Järvinen *et al.*)

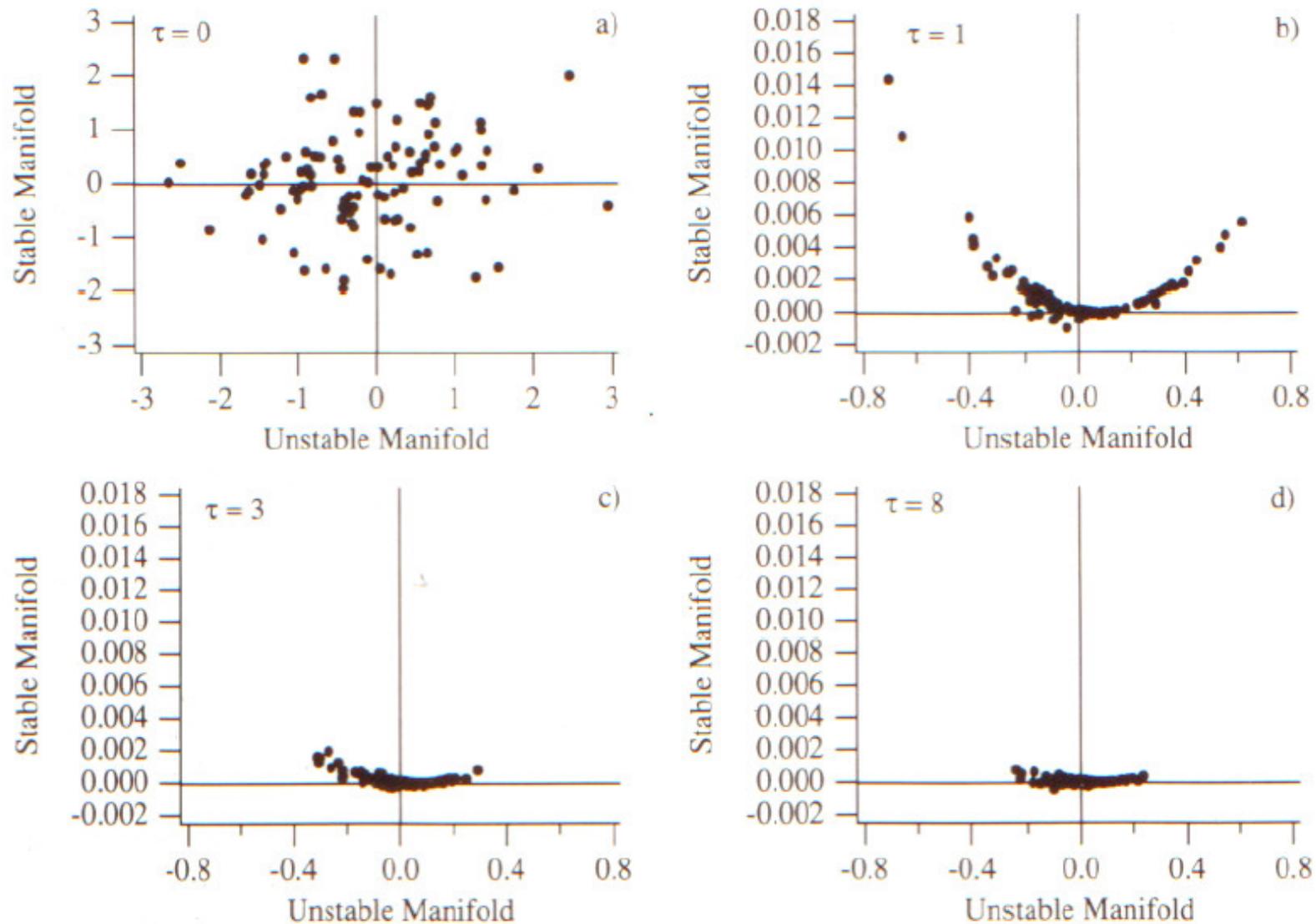


Fig. 7. Projection of the 100 minimizing solutions, at the end of the assimilation period, onto the plane spanned by the stable and unstable directions, defined as in Fig. 3. Values of τ are indicated on the panels. The projection is not an orthogonal projection, but a projection parallel to the local velocity vector $(dx/dt, dy/dt, dz/dt)$ (central manifold).

| $\mu(C(\tau, x))$ | Cloud of points QSVA | Cloud of points raw assimilation | Linear tangent system | Upper bound |
|-------------------|----------------------|----------------------------------|-----------------------|-------------|
| $\tau = 0$ | 1 | 1 | 1 | 1 |
| $\tau = 1$ | 0.36 | 0.37 | 0.39 | 0.46 |
| $\tau = 2$ | 5.9×10^{-2} | 5.74 | 4.5×10^{-2} | 0.401 |
| $\tau = 3$ | 3.3×10^{-2} | 29.4 | 2.9×10^{-2} | 0.397 |
| $\tau = 8$ | 1.4×10^{-2} | 59.9 | * | 0.396 |

In the left column, the estimates are calculated from the ensemble of 100 assimilations (see also Fig. 7). The 2nd column contains the values obtained from the raw assimilation. In the 3rd column, the estimates are obtained from the tangent linear system and eqs. (3.5–3.9) (the star indicates a computational overflow). The estimates in the right-hand column are the upper bounds defined by eq. (3.13).

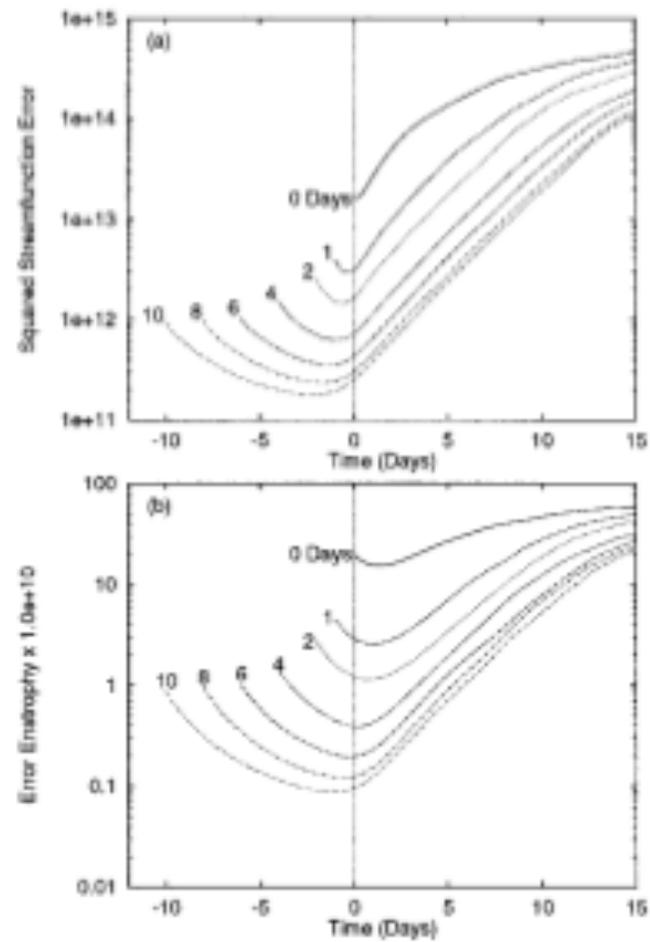


Fig. 5. Median values of the (a) streamfunction squared error, and (b) enstrophy error for the 200 forecast set as a function of forecast time and of the assimilation time T_a .

Swanson, Vautard and Pires, 1998, *Tellus*, **50A**, 369-390

Since, after an assimilation has been performed over a period of time, uncertainty is likely to be concentrated in modes that have been unstable, it might be useful for the next assimilation, and at least in terms of cost efficiency, to concentrate corrections on the background in those modes.

Actually, presence of residual noise in stable modes can be damageable for analysis and subsequent forecast.

Assimilation in the Unstable Subspace (AUS) (Carrassi *et al.*, 2007, 2008, for the case of 3D-Var)

Four-dimensional variational assimilation in the unstable subspace
(4DVar-AUS)

Trevisan *et al.*, 2010, Four-dimensional variational assimilation in the unstable subspace and the optimal subspace dimension, *Q. J. R. Meteorol. Soc.*, **136**, 487-496.

4D-Var-AUS

Algorithmic implementation

Define N perturbations to the current state, and evolve them according to the tangent linear model, with periodic reorthonormalization in order to avoid collapse onto the dominant Lyapunov vector (same algorithm as for computation of Lyapunov exponents).

Cycle successive 4D-Var's, restricting at each cycle the modification to be made on the current state to the space spanned by the N perturbations emanating from the previous cycle (if N is the dimension of state space, that is identical with standard 4D-Var).

Experiments performed on the Lorenz (1996) model

$$\frac{d}{dt}x_j = (x_{j+1} - x_{j-2})x_{j-1} - x_j + F$$

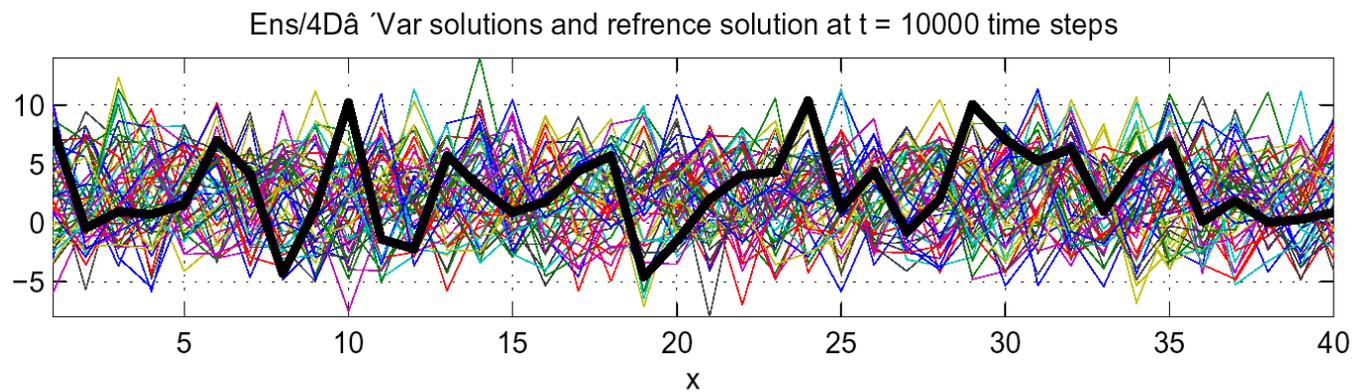
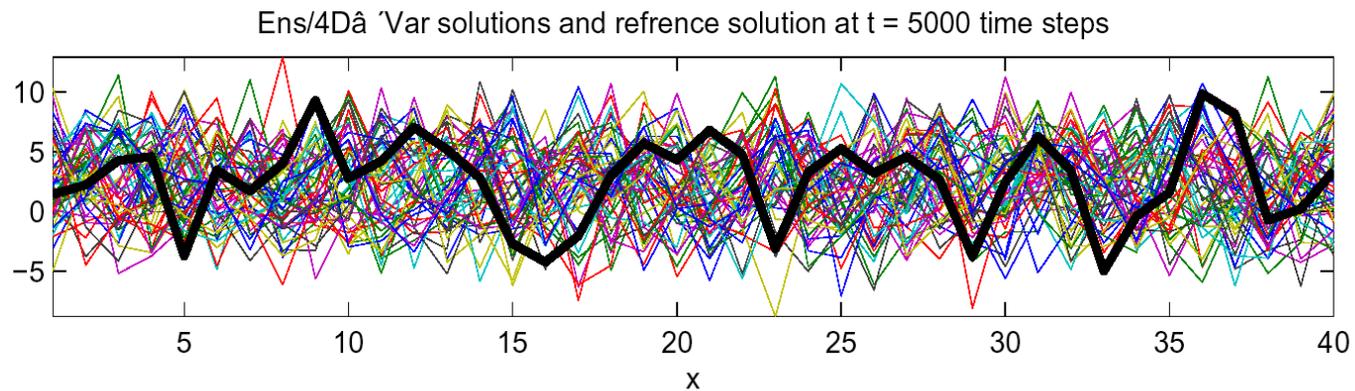
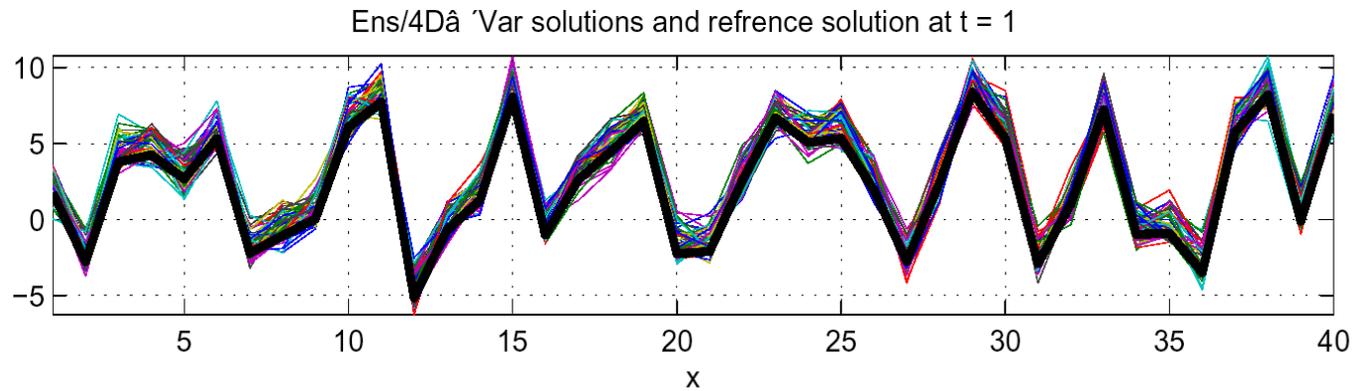
with $j = 1, \dots, I$.

with value $F = 8$, which gives rise to chaos.

Three values of I have been used, namely $I = 40, 60, 80$, which correspond to respectively $N^+ = 13, 19$ and 26 positive Lyapunov exponents.

In all three cases, the largest Lyapunov exponent corresponds to a doubling time of about 2 days (with 1 'day' = 1/5 model time unit).

Identical twin experiments (perfect model)



Lorenz'96 model (M. Jardak)

‘Observing system’ defined as in Fertig *et al.* (*Tellus*, 2007):

At each observation time, one observation every four grid points (observation points shifted by one grid point at each observation time).

Observation frequency : 1.5 hour

Random gaussian observation errors with expectation 0 and standard deviation $\sigma_0 = 0.2$ (‘climatological’ standard deviation 5.1).

Sequences of variational assimilations have been cycled over windows with length $\tau = 1, \dots, 5$ days. Results are averaged over 5000 successive windows.

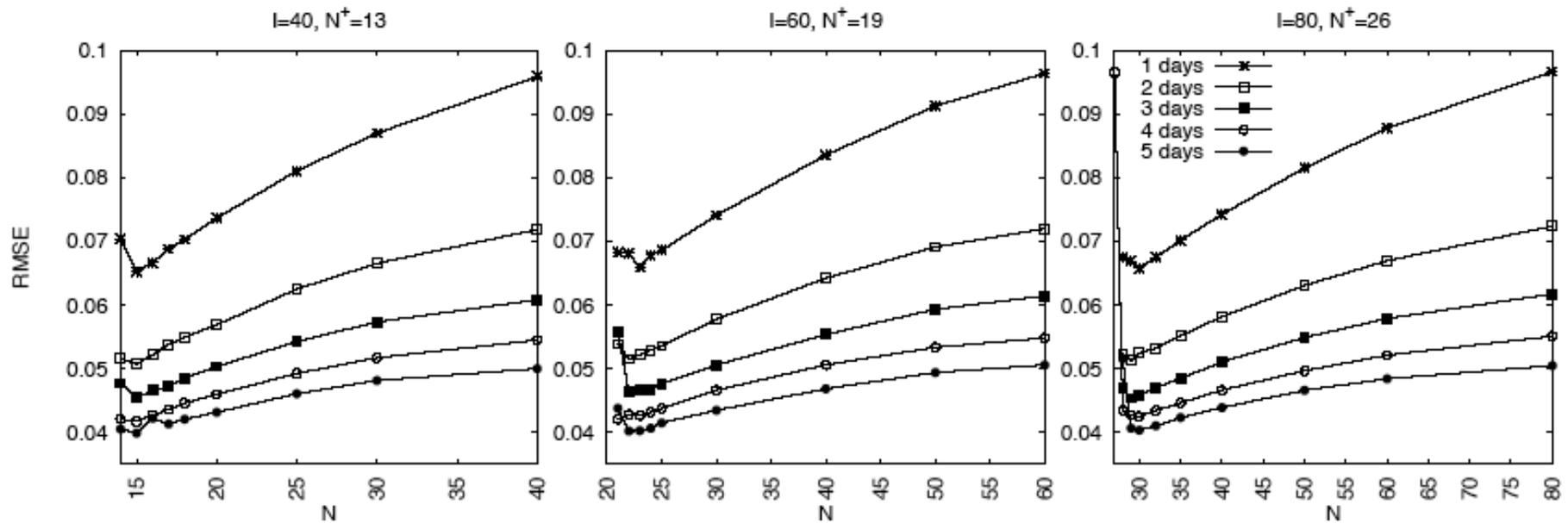


Figure 1. Time average RMS analysis error at $t = \tau$ as a function of the subspace dimension N for three model configurations: $I=40, 60, 80$. Different curves in the same panel refer to different assimilation windows from 1 to 5 days. The observation error standard deviation is $\sigma_o = 0.2$.

No explicit background term (*i. e.*, with error covariance matrix) in objective function : information from past lies in the background to be updated, and in the N perturbations which define the subspace in which updating is to be made.

Best performance for N slightly above number N^+ of positive Lyapunov exponents.

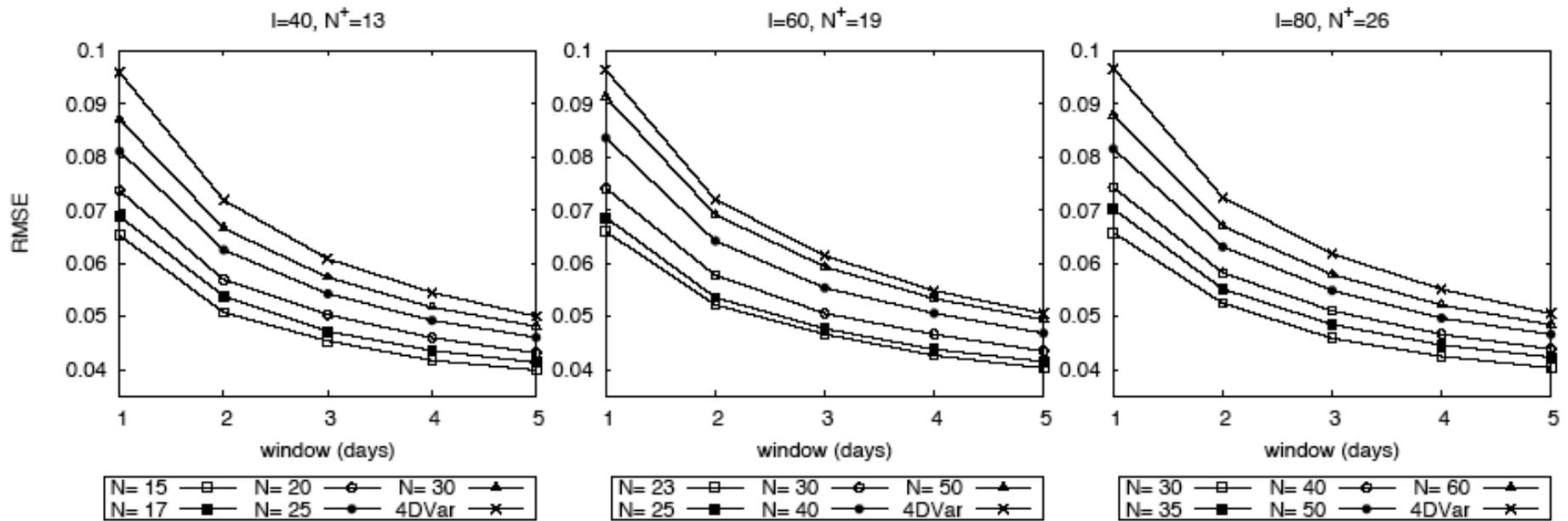


Figure 2. Time average RMS analysis error at $t = \tau$ as a function of the length of the assimilation window for three model configurations: $I=40, 60, 80$. Different curves in the same panel refer to a different subspace dimension N of 4DVar-AUS and to standard 4DVar. $\sigma_o = 0.2$.

Different curves are almost identical on all three panels. Relative improvement obtained by decreasing subspace dimension N to its optimal value is largest for smaller window length τ .

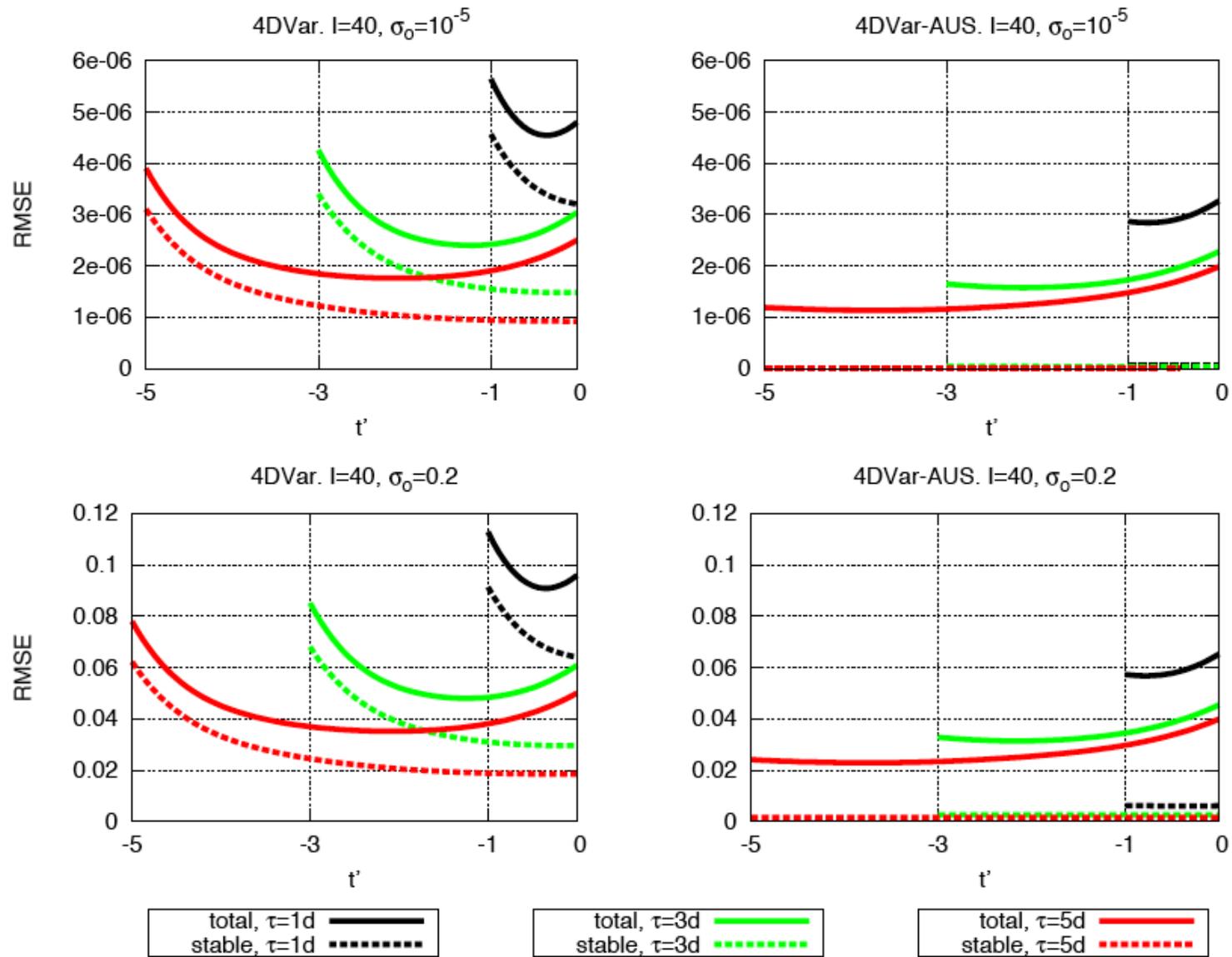


Figure 3. Time average RMS error within 1, 3, 5 days assimilation windows as a function of $t' = t - \tau$, with $\sigma_0 = .2, 10^{-5}$ for the model configuration $I = 40$. Left panel: 4DVar. Right panel: 4DVar-AUS with $N = 15$. Solid lines refer to total assimilation error, dashed lines refer to the error component in the stable subspace e_{16}, \dots, e_{40} .

Experiments have been performed in which an explicit background term was present, the associated error covariance matrix having been obtained as the average of a sequence of full **4D-Var**'s.

The estimates are systematically improved, and more for full **4D-Var** than for **4D-Var-AUS**. But they remain qualitatively similar, with best performance for **4D-Var-AUS** with N slightly above N^+ .

Minimum of objective function cannot be made smaller by reducing control space. Numerical tests show that minimum of objective function is smaller (by a few percent) for full **4D-Var** than for **4D-Var-AUS**. Full **4D-Var** is closer to the noisy observations, but farther away from the truth. And tests also show that full **4D-Var** performs best when observations are perfect (no noise).

Results show that, if all degrees of freedom that are available to the model are used, the minimization process introduces components along the stable modes of the system, in which no error is present, in order to ensure a closer fit to the observations. This degrades the closeness of the fit to reality. The optimal choice is to restrict the assimilation to the unstable modes.

Can have major practical algorithmic implications.

Questions.

- Degree of generality of results ?

- Impact of model errors ?

Conclusions

- The main two classes of algorithms that presently exist for operational assimilation of observations in geophysical applications are Ensemble Kalman Filter (EnKF) and Variational Assimilation (4D-Var). They both are more or less empirical extensions to mildly nonlinear and nongaussian situations of algorithms which achieve Bayesian estimation in linear and gaussian situations.
- These two classes of algorithms produce useful results (and of comparable quality).
- They are far from optimality. There is no obvious way for improving on them, but ensemble methods, meant to produce a sample of the sought-for conditional probability distribution, seem most promising. Fully bayesian particle filters are the subject of active research.

Conclusions (continued)

Assimilation, which originated from the need of defining initial conditions for numerical weather forecasts, has progressively extended to many diverse applications

- Oceanography
- Atmospheric chemistry (both troposphere and stratosphere)
- Oceanic biogeochemistry
- Ground hydrology
- Terrestrial biosphere and vegetation cover
- Glaciology
- Magnetism (both planetary and stellar)
- Plate tectonics
- Planetary atmospheres (Mars, ...)
- Reassimilation of past observations (mostly for climatological purposes, ECMWF, NCEP/NCAR)
- Identification of source of tracers
- Parameter identification
- *A priori* evaluation of anticipated new instruments
- Definition of observing systems (*Observing Systems Simulation Experiments*)
- Validation of models
- Sensitivity studies (adjoints)
- ...

ECMWF

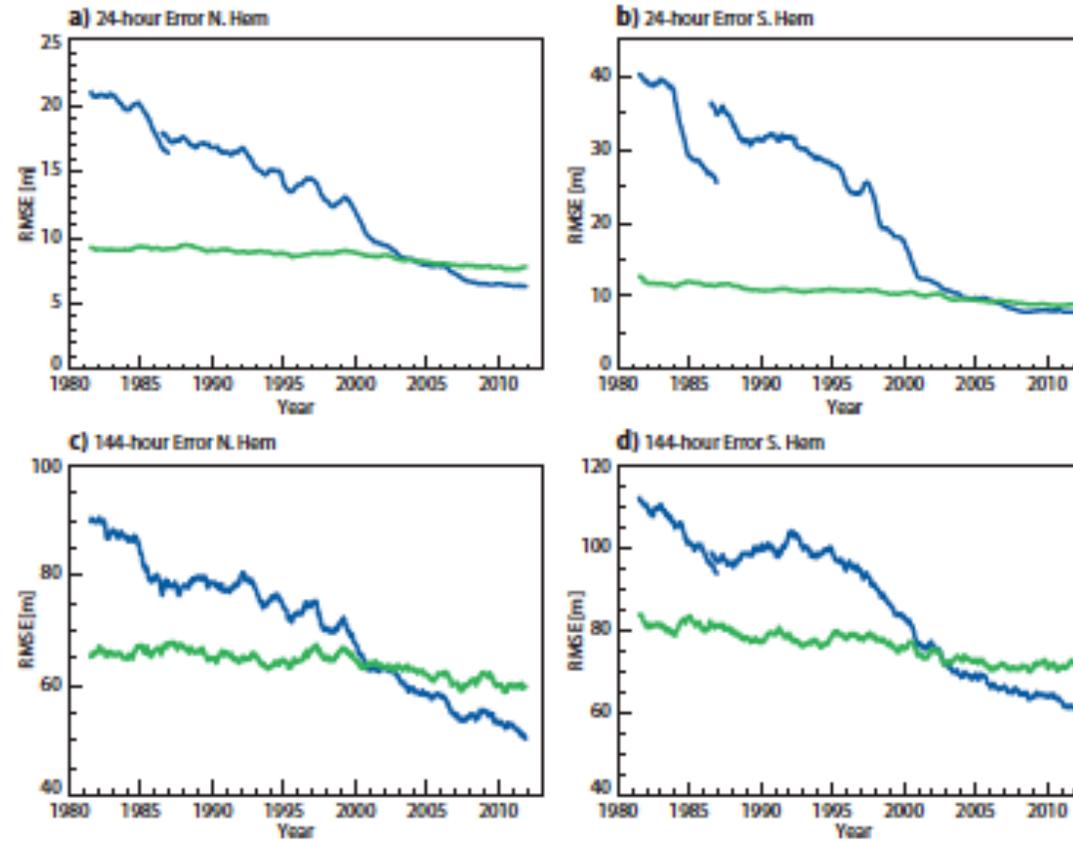


FIG. 3. Evolution of forecast errors from 1981 to 2012 for N.Hem (a and c) and S.Hem (b and d). Operational forecasts (blue) and ERA Interim (green). Note that before 1986 the operational analysis is used to verify the operational forecasts, after 1986 ERA Interim is used for the verification (with an overlap of 6 months present).

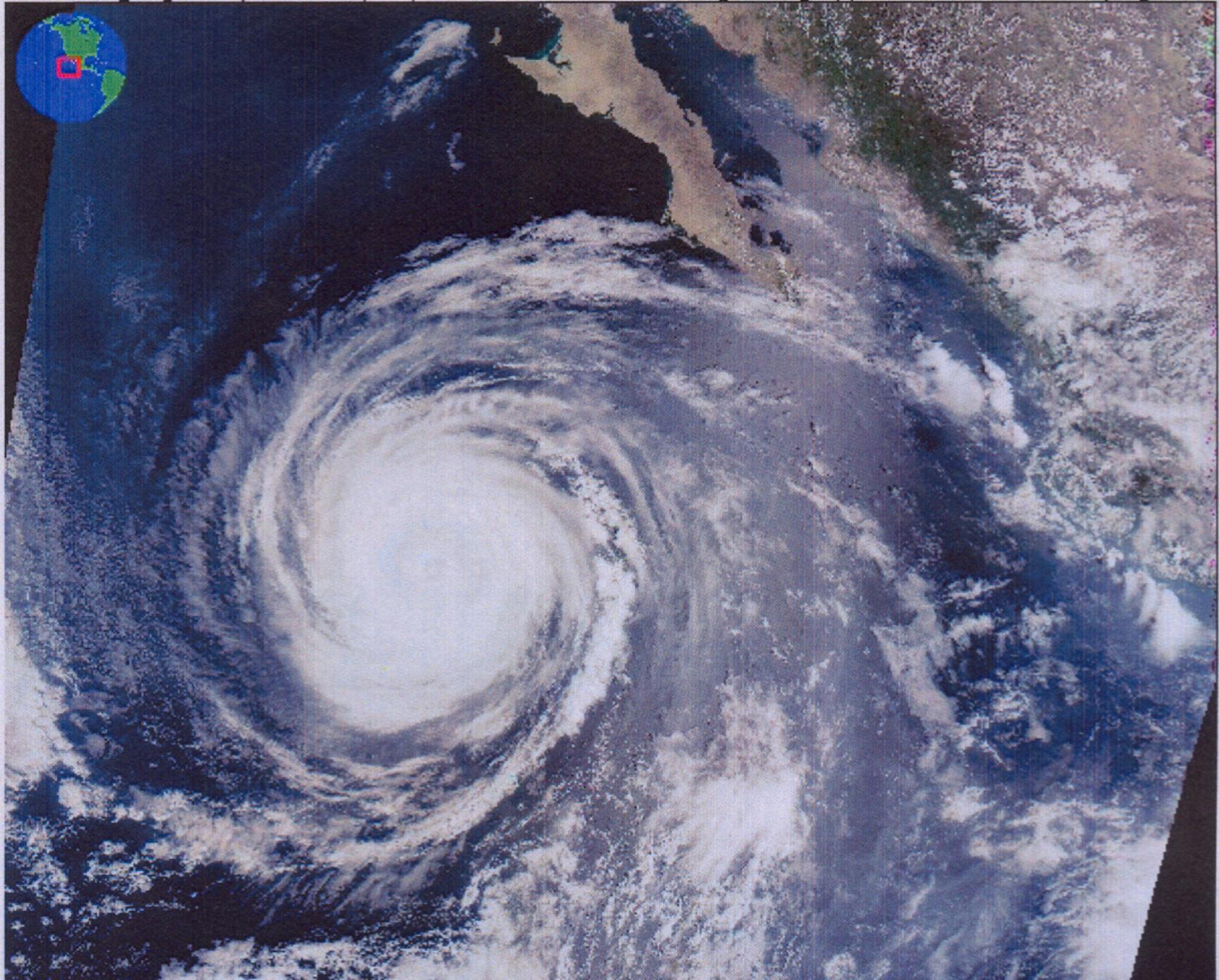
Assimilation is related to

- Estimation theory
- Probability theory
- Atmospheric and oceanic dynamics
- Atmospheric and oceanic predictability
- Instrumental physics
- Optimisation theory
- Control theory
- Algorithmics and computer science
- ...

A few of the (many) remaining problems (in addition to problems raised by practical numerical implementation):

- Observability (data are noisy, system is chaotic !)
- More accurate identification and quantification of errors affecting data particularly the assimilating model (will always require independent hypotheses)
- Assimilation of images
- ...

. HDFLook project (LOA-USTL) (MODIS October 2 2002 [18h10] ((Hurricane Hernan (Baja Cali

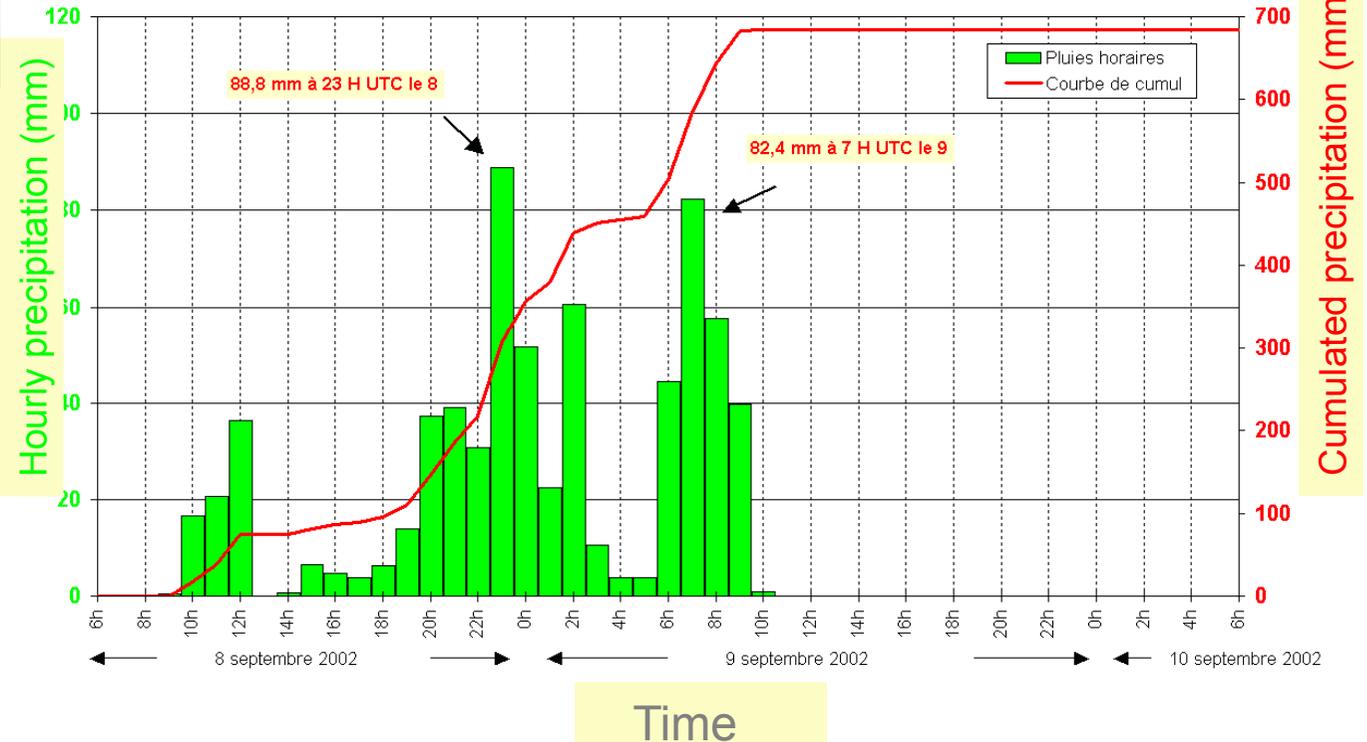




8-9 September 2002 :
More than 20 deaths
Economic loss : 1,6 billion €

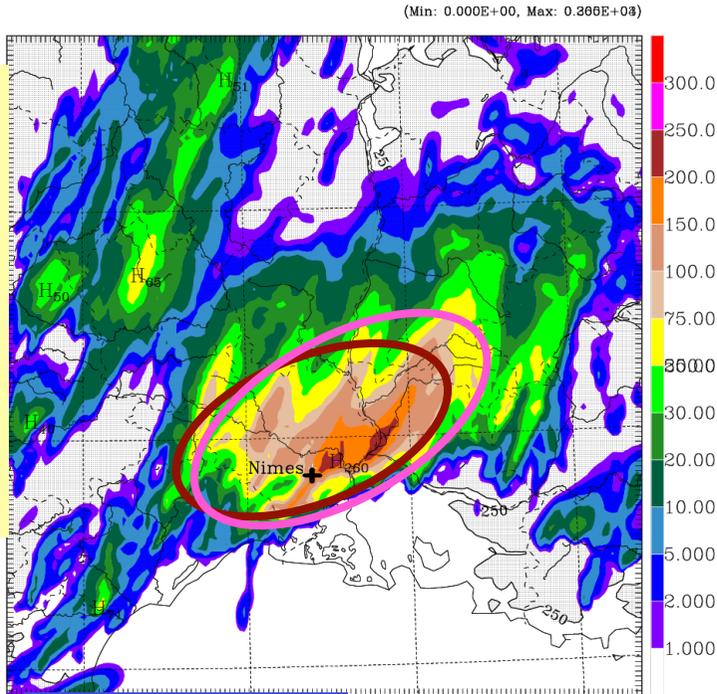


Précipitations à Cardet (alt: 115 m, Gard)
du 8/09/2002 à 6h (UTC) au 10/09/2002 à 6h (UTC)



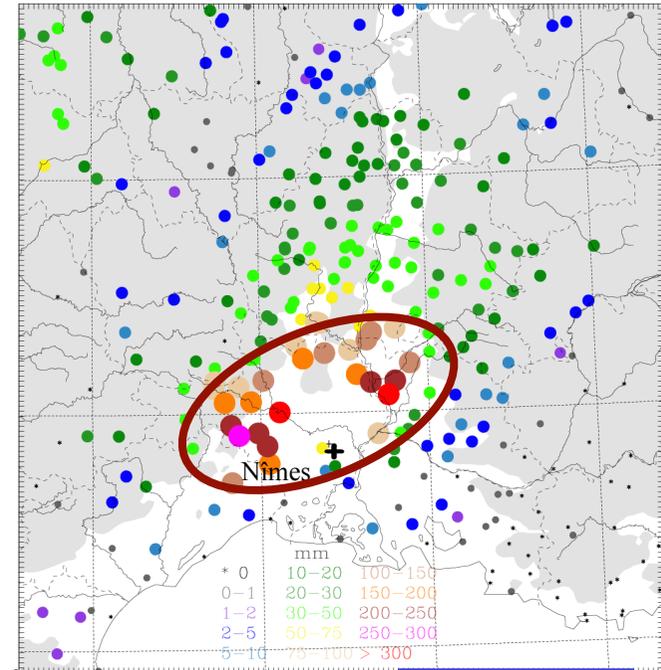
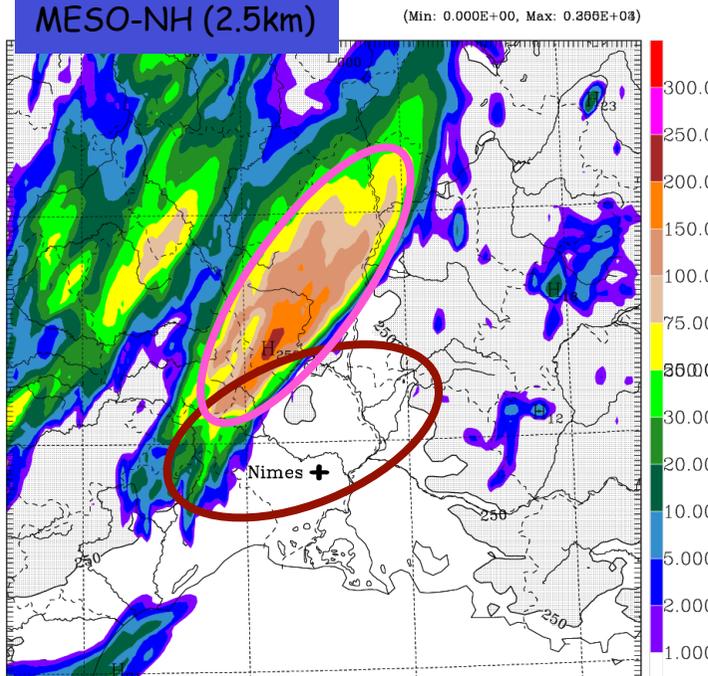
The accumulated rainfall totals have locally reached 600-700 mm

Initial conditions : mesoscale analysis (surface obs, radar, satellite) for 12UTC, 8th Sept. 2002



MESO-NH (2.5km)

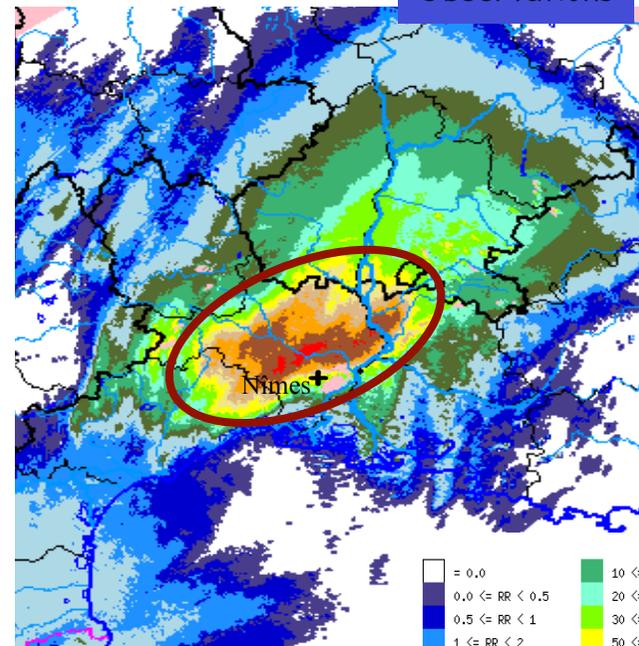
Initial conditions : large scale ARPEGE analysis for 12UTC, 8th Sept. 2002



Raingauges

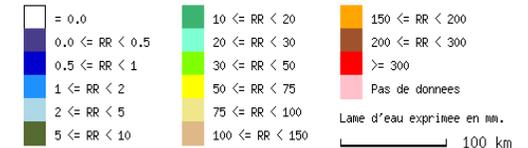
Observations

Ducrocq et al, 2003

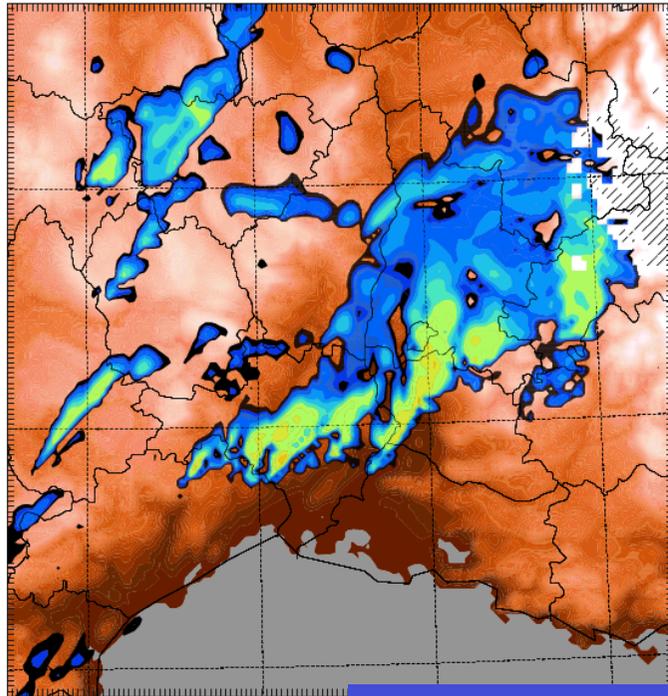


Nîmes radar

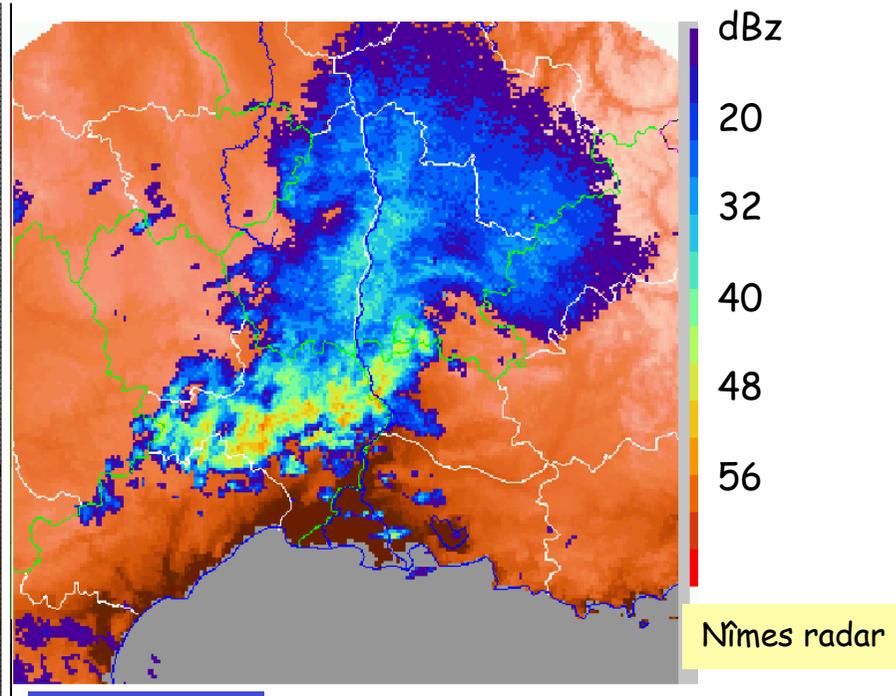
12-h accumulated rainfall from 12 UTC, 8 Sept to 0 UTC, 9 Sept 2002



Initial conditions : mesoscale analysis (surface obs, radar, satellite) for 12UTC, 8th Sept. 2002



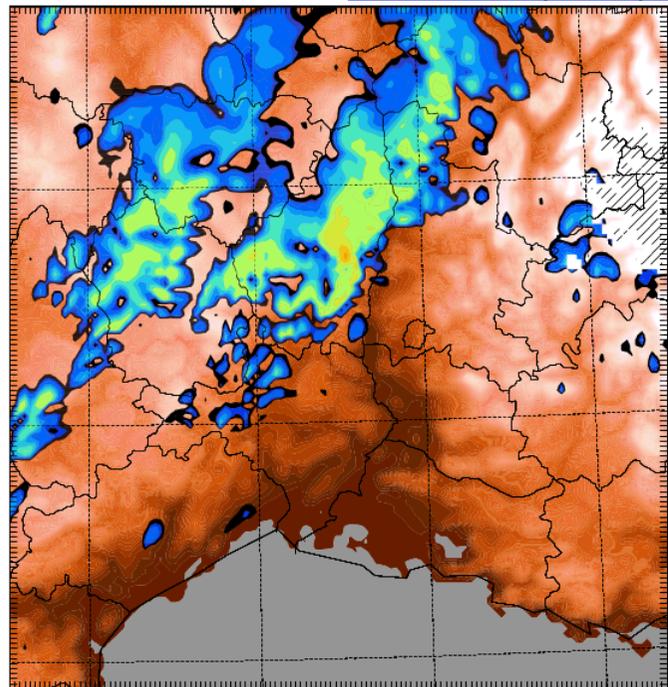
MESO-NH (2.5km)



Nîmes radar

Observations

Initial conditions : large scale ARPEGE analysis for 12UTC, 8th Sept. 2002



Reflectivities at 21 UTC, 8 Sept.

- **Références de base sur l'assimilation des observations**

- (O. Talagrand, Mai 2013)

-

- **Livres**

-

- *Problèmes Inverses en général*

-

- Tarantola, A., 2005, *Inverse Problem Theory and Methods for Model Parameter Estimation*, Society for Industrial and Applied Mathematics, Philadelphie, États-Unis, 342 pp., ISBN 0-89871-572-5 (peut-être téléchargé gratuitement à l'adresse <http://www.ipgp.jussieu.fr/~tarantola/>).

-

-

- *Théorie du Filtrage*

-

- Jazwinski, A. H., 1970, *Stochastic Processes and Filtering Theory*, Academic Press, New York, États-Unis, 376 pp..

-

- Gelb, A., 1974, *Applied optimal estimation*, Cambridge University Press, Cambridge, Royaume-Uni, 374 pp..

-

- Anderson, B. D. O., and J. B. Moore, 1979, *Optimal Filtering*, Prentice Hall, Englewood Cliffs, New Jersey, États-Unis, 357 pp..

- **Références de base sur l'assimilation des observations (suite)**
-
- *Applications Météorologiques et Océanographiques*
-
- Daley, R., 1991, *Atmospheric Data Analysis*, Cambridge University Press, Cambridge, Royaume-Uni, 457 pp..
-
- Bennett, A. F., 1992, *Inverse Methods in Physical Oceanography*, Cambridge University Press, Cambridge, Royaume-Uni, 346 pp..
-
- Wunsch, C., 1996, *The Ocean Circulation as an Inverse Problem*, Cambridge University Press, Cambridge, Royaume-Uni, 442 pp..
-
- Rodgers, C. D., 2000, *Inverse Methods for Atmospheric Sounding: Theory and Practice*, World Scientific Publishing Co. Ltd, Londres, Royaume-Uni, 238 pp..
-
- Kalnay, E., 2002, *Atmospheric Modeling, Data Assimilation and Predictability*, Cambridge University Press, Cambridge, Royaume-Uni, 341 pp..
-
- (Bennett, A. F., 2002, *Inverse Modeling of the Ocean and Atmosphere*, Cambridge University Press, Cambridge, Royaume-Uni, 265 pp.).
-
- Evensen, G., 2009, *Data assimilation, The Ensemble Kalman Filter*, 2nd ed., Springer, 307 pp..

- **Références de base sur l'assimilation des observations (suite)**

-

-

- **Articles**

-

- Ghil, M., et P. Malanotte-Rizzoli, 1991, Data assimilation in meteorology and oceanography, *Adv. in Geophys.*, **33**, 141-266.

-

- Talagrand, O., 1997, Assimilation of Observations, an Introduction, *J. Meteor. Soc. Japan*, **75** (1B, Numéro spécial Data Assimilation in Meteorology and Oceanography: Theory and Practice), 191-20 (et autres articles dans le même numéro spécial).