

Mathematics/Hydrodynamics/Geophysical Fluid Dynamics Refresher Course

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M1 MOCIS

Necessary mathematics

Vector algebra
Differential operations on
scalar and vector fields
Integration in 3D space
Fourier analysis

Basic notions of wave dynamics

Simple-wave equation
Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid
Governing equations
Euler - Lagrange duality
Energy and
thermodynamics
Kelvin circulation theorem
Real fluids: incorporating
molecular transport

Hydrodynamics on a tangent plane to a rotating planet

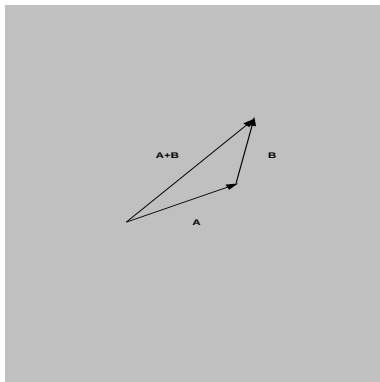
Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean
Primitive equations:
Atmosphere
Equations in pressure
coordinates
Dynamical approximations

Vectors: definitions and superposition principle

Vector \mathbf{A} is a coordinate-independent (invariant) object having a magnitude $|\mathbf{A}|$ and a direction. Alternative notation \vec{A} .

Adding/subtracting vectors:



Superposition principle: Linear combination of vectors is a vector

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Products of vectors

Scalar product of two vectors:

Projection of one vector onto another:

$$\mathbf{A} \cdot \mathbf{B} := |\mathbf{A}| |\mathbf{B}| \cos \phi_{AB} \equiv \mathbf{B} \cdot \mathbf{A},$$

where ϕ_{AB} is an included angle between the two.

Vector product of two vectors:

$$\mathbf{A} \wedge \mathbf{B} := \hat{\mathbf{i}}_{AB} |\mathbf{A}| |\mathbf{B}| \sin \phi_{AB} = -\mathbf{B} \wedge \mathbf{A},$$

where $\hat{\mathbf{i}}_{AB}$ is a unit vector, $|\hat{\mathbf{i}}_{AB}| = 1$, perpendicular to both \mathbf{A} and \mathbf{B} , with the orientation of a right-handed screw rotated from \mathbf{A} toward \mathbf{B} .

\times is an alternative notation for \wedge .

Distributive properties:

$$(\mathbf{A} + \mathbf{B}) \cdot \mathbf{C} = \mathbf{A} \cdot \mathbf{C} + \mathbf{B} \cdot \mathbf{C}, \quad (\mathbf{A} + \mathbf{B}) \wedge \mathbf{C} = \mathbf{A} \wedge \mathbf{C} + \mathbf{B} \wedge \mathbf{C}.$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields
Integration in 3D space
Fourier analysis

Basic notions of wave dynamics

Simple-wave equation
Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

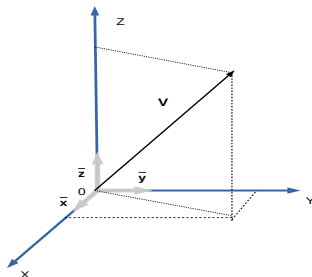
Governing equations
Euler - Lagrange duality
Energy and thermodynamics
Kelvin circulation theorem
Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean
Primitive equations: Atmosphere
Equations in pressure coordinates
Dynamical approximations

Vectors in Cartesian coordinates



Cartesian coordinates: defined by a right triad of mutually orthogonal unit vectors forming a **basis**:

$$(\hat{x}, \hat{y}, \hat{z}) \equiv (\hat{x}_1, \hat{x}_2, \hat{x}_3),$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields
Integration in 3D space
Fourier analysis

Basic notions of wave dynamics

Simple-wave equation
Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations
Euler - Lagrange duality
Energy and thermodynamics
Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean
Primitive equations: Atmosphere
Equations in pressure coordinates
Dynamical approximations

Tensor notation and Kronecker delta

$(\hat{\mathbf{x}}, \hat{\mathbf{y}}, \hat{\mathbf{z}}) \rightarrow \hat{\mathbf{x}}_i, i = 1, 2, 3$. Ortho-normality of the basis:

$$\hat{\mathbf{x}}_i \cdot \hat{\mathbf{x}}_j = \delta_{ij},$$

where δ_{ij} is Kronecker delta-symbol, an invariant **tensor** of second rank (3×3 unit diagonal matrix):

$$\delta_{ij} = \begin{cases} 1, & \text{if } i = j, \\ 0, & \text{if } i \neq j. \end{cases}$$

The components V_i of a vector \mathbf{V} are given by its *projections* on the axes $V_i = \mathbf{V} \cdot \hat{\mathbf{x}}_i$:

$$\mathbf{V} = V_1 \hat{\mathbf{x}}_1 + V_2 \hat{\mathbf{x}}_2 + V_3 \hat{\mathbf{x}}_3 \equiv \sum_{i=1}^3 V_i \hat{\mathbf{x}}_i$$

Einstein's convention:

$\sum_{i=1}^3 A_i B_i \equiv A_i B_i$ (self-repeating index is “dumb”).

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields
Integration in 3D space
Fourier analysis

Basic notions of wave dynamics

Simple-wave equation
Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations
Euler - Lagrange duality
Energy and thermodynamics
Kelvin circulation theorem
Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean
Primitive equations: Atmosphere
Equations in pressure coordinates
Developing approximations

Vector products by Levi-Civita tensor

Formula for the vector product:

$$\mathbf{A} \wedge \mathbf{B} = \begin{vmatrix} \hat{x} & \hat{y} & \hat{z} \\ A_1 & A_2 & A_3 \\ B_1 & B_2 & B_3 \end{vmatrix}$$

Tensor notation (with Einstein's convention):

$$(\mathbf{A} \wedge \mathbf{B})_i = \epsilon_{ijk} A_j B_k,$$

where

$$\epsilon_{ijk} = \begin{cases} 1, & \text{if } ijk = 123, 231, 312 \\ -1, & \text{if } ijk = 132, 321, 213 \\ 0, & \text{otherwise} \end{cases}$$

Magic identity:

$$\epsilon_{ijk} \epsilon_{klm} = \delta_{il} \delta_{jm} - \delta_{im} \delta_{jl}. \quad (1)$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Scalar, vector, and tensor fields

Any point in space is given by its **radius-vector**

$$\mathbf{x} = x\hat{\mathbf{x}} + y\hat{\mathbf{y}} + z\hat{\mathbf{z}}.$$

A **field** is an object defined at any point of space

$(x, y, z) \equiv (x_1, x_2, x_3)$ at any moment of time t , i.e. a function of \mathbf{x} and t .

Different types of fields:

- ▶ scalar $f(\mathbf{x}, t)$,
- ▶ vector $\mathbf{v}(\mathbf{x}, t)$,
- ▶ tensor $t_{ij}(\mathbf{x}, t)$

The fields are **dependent variables**, and x, y, z and t - **independent variables**.

Physical examples: scalar fields - temperature, density, pressure, geopotential, vector fields - velocity, electric and magnetic fields, tensor fields - stresses, gravitational field.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Developing approximations

Differential operations on scalar fields

Partial derivatives:

$$\frac{\partial f}{\partial x} := \lim_{\Delta x \rightarrow 0} \frac{f(x + \Delta x, y, z) - f(x, y, z)}{\Delta x},$$

and similar for other independent variables. Differential operator **nabla**:

$$\nabla := \hat{\mathbf{x}} \frac{\partial}{\partial x} + \hat{\mathbf{y}} \frac{\partial}{\partial y} + \hat{\mathbf{z}} \frac{\partial}{\partial z}$$

Gradient of a scalar field: the vector field

$$\text{grad } f \equiv \nabla f = \hat{\mathbf{x}} \frac{\partial f}{\partial x} + \hat{\mathbf{y}} \frac{\partial f}{\partial y} + \hat{\mathbf{z}} \frac{\partial f}{\partial z},$$

Heuristic meaning: a vector giving direction and rate of fastest increase of the function f .

Necessary
mathematics

Vector algebra

Differential operations on
scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of
wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in
fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and
thermodynamics

Kelvin circulation theorem

Real fluids: incorporating
molecular transport

Hydrodynamics on
a tangent plane to
a rotating planet

Primitive equations
for the ocean and
the atmosphere

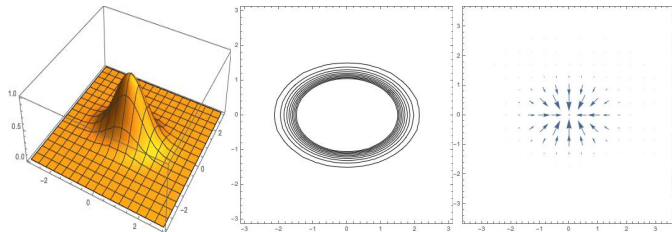
Primitive equations: Ocean

Primitive equations:
Atmosphere

Equations in pressure
coordinates

Equations in isopycnal

Visualizing gradient in 2D



From left to right: 2D relief, its contour map, and its gradient. Graphics by Mathematica[©]

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Differential operations with vectors

- ▶ Scalar product: divergence

$$\operatorname{div} \mathbf{v} \equiv \nabla \cdot \mathbf{v}(\mathbf{x}) = \frac{\partial v_i}{\partial x_i}$$

- ▶ Vector product: curl

$$\operatorname{curl} \mathbf{v} \equiv \nabla \wedge \mathbf{v}(\mathbf{x}); \quad (\operatorname{curl} \mathbf{v})_i = \epsilon_{ijk} \frac{\partial v_k}{\partial x_j}$$

- ▶ Tensor product:

$$\nabla \otimes \mathbf{v}(\mathbf{x}); \quad (\nabla \otimes \mathbf{v})_{ij} = \frac{\partial v_i}{\partial x_j}$$

For any \mathbf{v} , f : $\operatorname{div} \operatorname{curl} \mathbf{v} \equiv 0$, $\operatorname{curl} \operatorname{grad} f \equiv 0$,
 $\operatorname{div} \operatorname{grad} f = \nabla^2 f$, $\nabla^2 = \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}$ - **Laplacian**.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

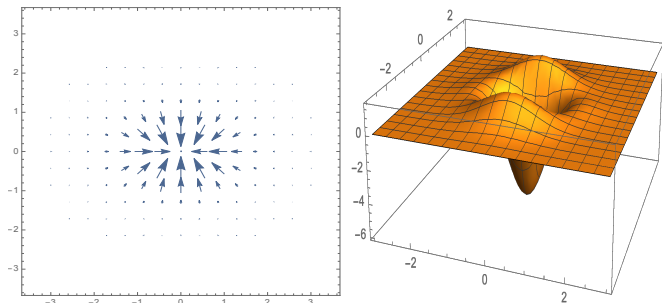
Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Visualizing divergence in 2D



From left to right: vector field $\mathbf{v}(x, y) = (v_1(x, y), v_2(x, y))$, and its divergence $\frac{\partial v_1}{\partial x} + \frac{\partial v_2}{\partial y}$. The curl $\hat{\mathbf{z}} \left(\frac{\partial v_2}{\partial x} - \frac{\partial v_1}{\partial y} \right)$ of this field is identically zero. (The field is a gradient of the previous example.) Graphics by Mathematica[©]

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

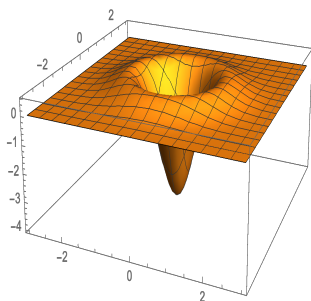
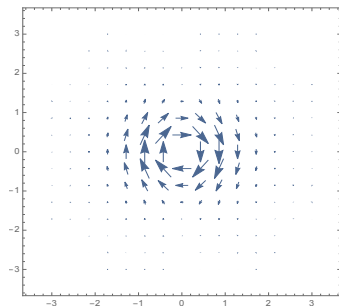
Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Developing approximations

Visualizing curl in 2D



From left to right: vector field $\mathbf{v}(x, y) = (v_1(x, y), v_2(x, y))$, and its curl $\frac{\partial v_2}{\partial x} - \frac{\partial v_1}{\partial y}$. The divergence $\frac{\partial v_1}{\partial x} + \frac{\partial v_2}{\partial y}$ of this field is identically zero, so the field is a curl of another vector field. Graphics by Mathematica[©]

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

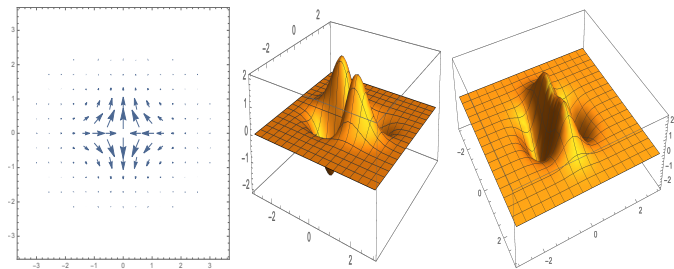
Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Strain field with non-zero curl and divergence



From left to right: vector field, and its curl and divergence.

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Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Developing approximations

Useful identities

$$\nabla \wedge (\nabla \wedge \mathbf{v}) = \nabla(\nabla \cdot \mathbf{v}) - \nabla^2 \mathbf{v}, \quad (2)$$

$$\mathbf{v} \wedge (\nabla \wedge \mathbf{v}) = \nabla \left(\frac{\mathbf{v}^2}{2} \right) - (\mathbf{v} \cdot \nabla) \mathbf{v}, \quad (3)$$

$$\nabla f \cdot (\nabla \wedge \mathbf{v}) = -\nabla \cdot (\nabla f \wedge \mathbf{v}). \quad (4)$$

Proofs: using tensor representation $(\nabla \wedge \mathbf{v})_i = \epsilon_{ijk} \partial_j v_k$, with shorthand notation $\frac{\partial}{\partial x_i} \equiv \partial_i$, exploiting the antisymmetry of ϵ_{ijk} , using that $\delta_{ij} v_j = v_i$, and applying the magic formula (1).

Example: proof of (2).

$$\epsilon_{ijk} \partial_j \epsilon_{klm} \partial_l v_m = (\delta_{il} \delta_{jm} - \delta_{im} \delta_{jl}) \partial_j \partial_l v_m = \partial_i \partial_j v_j - \partial_j \partial_j v_i.$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

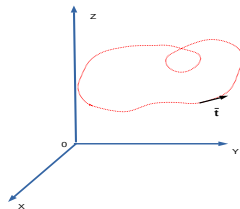
Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Vertical coordinate systems

Integration of a field along a (closed) 1D contour



Summation of the values of the field at the points of the contour times oriented line element $d\mathbf{l} = \hat{\mathbf{t}} dl$:

$$\oint d\mathbf{l}(\dots),$$

where $\hat{\mathbf{t}}$ is unit tangent vector, and dl is a length element along the contour. Positive orientation: anti-clockwise.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

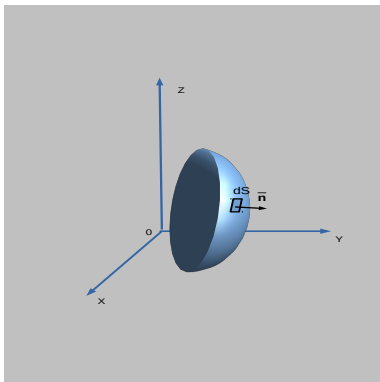
Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Vertical coordinate transformation

Integration of a field over a 2D surface



Summation of the values of the field at the points of the surface times oriented surface element $d\mathbf{s} = \hat{\mathbf{n}} ds$:

$$\iint d\mathbf{s}(\dots) \equiv \int_S d\mathbf{s}(\dots),$$

where $\hat{\mathbf{n}}$ is unit normal vector. Positive orientation for closed surfaces: outwards.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

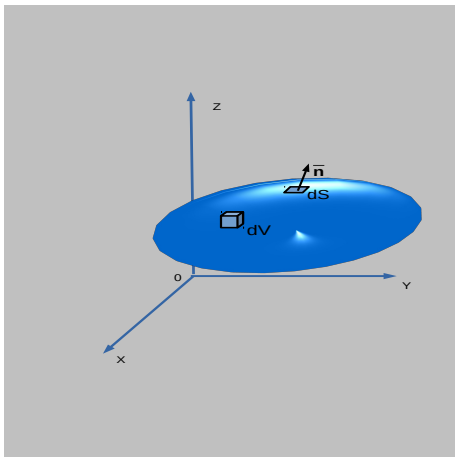
Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Integration of a field over a 3D volume



Summation of the values of the field at the points in the volume times volume element dV .

$$\iiint dV(\dots) \equiv \int_V dV(\dots).$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

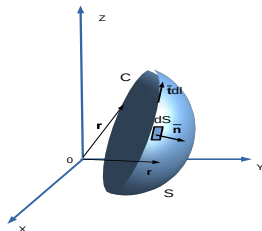
Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Vertical coordinate systems

Linking contour and surface integrations: Stokes theorem



$$\oint_C d\mathbf{l} \cdot \mathbf{v}(\mathbf{x}) = \int_{S_C} d\mathbf{s} \cdot (\nabla \wedge \mathbf{v}(\mathbf{x})). \quad (5)$$

Left-hand side: **circulation** of the vector field over the contour C . Right-hand side: curl of \mathbf{v} integrated over **any** surface S_C having the contour C as a base.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

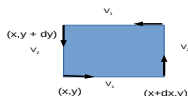
Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Stokes theorem: the idea of proof



Circulation of the vector $\mathbf{v} = v_1 \hat{\mathbf{x}} + v_2 \hat{\mathbf{y}}$ over an elementary contour, with $dx \rightarrow 0$, $dy \rightarrow 0$, using first-order Taylor expansions:

$$\begin{aligned} v_1(x, y)dx + v_2(x + dx, y)dy - v_1(x, y + dy)dx - v_2(x, y)dy \\ = \frac{\partial v_2}{\partial x} dx dy - \frac{\partial v_1}{\partial y} dx dy, \end{aligned}$$

with a z-component of $\text{curl} \mathbf{v}$ multiplied by the z-oriented surface element arising in the right-hand side.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Developing approximations

Linking surface and volume integrations: Gauss theorem

$$\oint_{S_V} d\mathbf{s} \cdot \mathbf{v}(\mathbf{x}) = \int_V dV \nabla \cdot \mathbf{v}(\mathbf{x}). \quad (6)$$

Left-hand side: **flux** of the vector field through the surface S_V which is a boundary of the volume V . Right-hand side: volume integral of the divergence of the field.

Important. The theorem is also valid for the scalar field:

$$\oint_{S_V} d\mathbf{s} \cdot \mathbf{f}(\mathbf{x}) = \int_V dV \nabla f(\mathbf{x}). \quad (7)$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

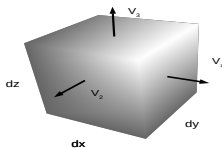
Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Gauss theorem: the idea of proof



Flux of the vector $\mathbf{v} = v_1 \hat{\mathbf{x}} + v_2 \hat{\mathbf{y}} + v_3 \hat{\mathbf{z}}$ over a surface of an elementary volume, taking into account the opposite orientation of the oriented surface elements:

$$\begin{aligned} & [v_1(x + dx, y, z) - v_1(x, y, z)] dydz + \\ & [v_2(x, y + dy, z) - v_2(x, y, z)] dx dz + \\ & [v_3(x, y, z + dz) - v_3(x, y, z)] dx dy = \left(\frac{\partial v_1}{\partial x} + \frac{\partial v_2}{\partial y} + \frac{\partial v_3}{\partial z} \right) dx dy dz \end{aligned}$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Fourier series for periodic functions

Consider $f(x) = f(x + 2\pi)$, a periodic smooth function on the interval $[0, 2\pi]$. **Fourier series:**

$$f(x) = \sum_{n=0}^{\infty} [a_n \cos(nx) + b_n \sin(nx)].$$

The expansion is unique due to **orthogonality** of the basis functions:

$$\int_0^{2\pi} dx \cos(nx) \cos(mx) = \int_0^{2\pi} dx \sin(nx) \sin(mx) = \pi \delta_{nm},$$

$$\int_0^{2\pi} dx \sin(nx) \cos(mx) \equiv 0.$$

The coefficients of expansion, thus, are uniquely defined:

$$a_n = \frac{1}{\pi} \int_0^{2\pi} dx f(x) \cos(nx), \quad b_n = \frac{1}{\pi} \int_0^{2\pi} dx f(x) \sin(nx)$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Complex exponential form

$$e^{inx} = \cos(nx) + i \sin(nx) \Rightarrow$$
$$\cos(nx) = \frac{e^{inx} + e^{-inx}}{2}, \quad \sin(nx) = \frac{e^{inx} - e^{-inx}}{2i}$$

Hence

$$f(x) = \sum_{n=0}^{\infty} \frac{(a_n - ib_n)}{2} e^{inx} + c.c \equiv \sum_{-\infty}^{\infty} A_n e^{inx}, \quad A_n^* = A_{-n}$$

Orthogonality:

$$\int_0^{2\pi} dx e^{inx} e^{-imx} = 2\pi \delta_{nm}$$

Expression for coefficients

$$A_n = \frac{1}{2\pi} \int_0^{2\pi} dx f(x) e^{-inx}$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Fourier integral

Fourier series on arbitrary interval L : $\sin(nx)$, $\cos(nx) \rightarrow \sin(\frac{2\pi}{L}nx)$, $\cos(\frac{2\pi}{L}nx)$, $\int_0^{2\pi} dx \rightarrow \int_0^L dx$, normalization $\frac{1}{\pi} \rightarrow \frac{1}{L}$. In the limit $L \rightarrow \infty$: $\sum_{-\infty}^{\infty} \rightarrow \int_{-\infty}^{\infty}$.

Fourier-transformation and its inverse:

$$f(x) = \int_{-\infty}^{\infty} dk F(k) e^{ikx}, \quad F(k) = \int_{-\infty}^{\infty} dx f(x) e^{-ikx}.$$

Based on orthogonality:

$$\int_{-\infty}^{\infty} dx e^{ikx} e^{-ilx} = \delta(k - l),$$

where $\delta(x)$ - Dirac's delta-function, continuous analog of Kronecker's δ_{nm} , with properties:

$$\int_{-\infty}^{\infty} dx \delta(x) = 1, \quad \int_{-\infty}^{\infty} dy \delta(x - y) F(y) = F(x).$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Vertical coordinate transformation

Multiple variables and differentiation

$$f(x, y, z) = \int_{-\infty}^{\infty} dk dl dm F(k, l, m) e^{i(kx+ly+mz)},$$

$$F(k, l, m) = \int_{-\infty}^{\infty} dx dy dz f(x, y, z) e^{-i(kx+ly+mz)}.$$

Physical space $(x, y, z) \rightarrow (k, l, m)$, Fourier space.

Radius-vector $\mathbf{x} \rightarrow \mathbf{k}$, "wavevector",

$$f(\mathbf{x}) = \int_{-\infty}^{\infty} d\mathbf{k} F(\mathbf{k}) e^{i\mathbf{k}\cdot\mathbf{x}}$$

Main advantage: differentiation in physical space \rightarrow multiplication by the corresponding component of the wavevector in Fourier space $\frac{\partial}{\partial x} \rightarrow ik$:

$$\frac{\partial}{\partial x} f(\mathbf{x}) = \int_{-\infty}^{\infty} d\mathbf{k} ik F(\mathbf{k}) e^{i\mathbf{k}\cdot\mathbf{x}},$$

and similarly for other variables.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Simplest wave equation

$$u_t + cu_x = 0. \quad (8)$$

$u(x, t)$ - dynamical variable, defined for all x :
 $-\infty < x < +\infty$, and t : $0 \leq t < \infty$, $c = \text{const.}$

Notation: $(\dots)_x = \frac{\partial(\dots)}{\partial x}$, $(\dots)_t = \frac{\partial(\dots)}{\partial t}$

Method of solution 1: change of variables.

$$(x, t) \rightarrow (\xi_+, \xi_-) = (x + ct, x - ct). \quad (9)$$

$$\frac{\partial \xi_{\pm}}{\partial x} = 1, \quad \frac{\partial \xi_{\pm}}{\partial t} = \pm c \Rightarrow \quad (10)$$

$$\frac{\partial u}{\partial t} = c \left(\frac{\partial u}{\partial \xi_+} - \frac{\partial u}{\partial \xi_-} \right), \quad (11)$$

$$\frac{\partial u}{\partial x} = \frac{\partial u}{\partial \xi_+} + \frac{\partial u}{\partial \xi_-} \quad (12)$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Simplification of the equation

$$u_t + cu_x = 0 \rightarrow 2c \frac{\partial u}{\partial \xi_+} = 0 \Rightarrow u = u(\xi_-). \quad (13)$$

Function u determined from initial conditions:

$$\text{c.l. : } u_{t=0} = u_0(x) \Rightarrow u = u_0(x - ct). \quad (14)$$

Necessary mathematics

- Vector algebra
- Differential operations on scalar and vector fields
- Integration in 3D space
- Fourier analysis

Basic notions of wave dynamics

- Simple-wave equation
- Dispersion, non-linearity

A crash course in fluid dynamics

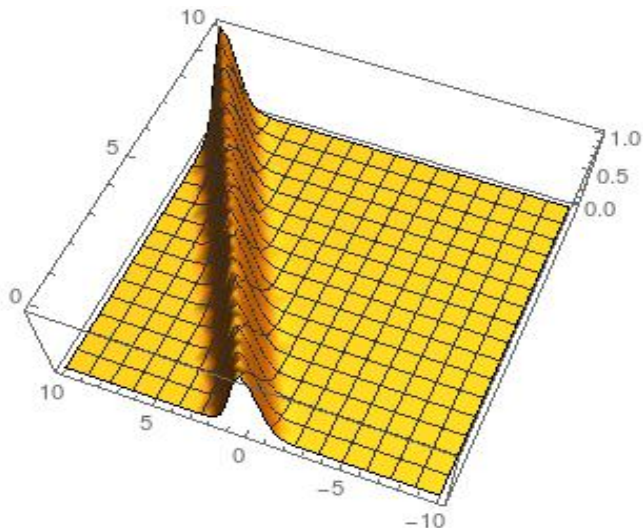
- The perfect fluid
 - Governing equations
 - Euler - Lagrange duality
 - Energy and thermodynamics
 - Kelvin circulation theorem
- Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

- Primitive equations: Ocean
- Primitive equations: Atmosphere
 - Equations in pressure coordinates
 - Vertical coordinate transformation

Spatio-temporal evolution of a localized initial perturbation, as follows from (8)



Necessary mathematics

- Vector algebra
- Differential operations on scalar and vector fields
- Integration in 3D space
- Fourier analysis

Basic notions of wave dynamics

- Simple-wave equation**
- Dispersion, non-linearity

A crash course in fluid dynamics

- The perfect fluid
 - Governing equations
 - Euler - Lagrange duality
 - Energy and thermodynamics
 - Kelvin circulation theorem
- Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

- Primitive equations: Ocean
- Primitive equations: Atmosphere
 - Equations in pressure coordinates
 - Vertical coordinate systems

Fourier transform

Method of solution 2: Fourier- transformation

$$u(x, t) = \frac{1}{2\pi} \int dk d\omega e^{i(kx - \omega t)} \hat{u}(k, \omega) + c.c.. \quad (15)$$

Inverse transformation:

$$\hat{u}(k, \omega) = \frac{1}{2\pi} \int dx dt e^{-i(kx - \omega t)} u(x, t) + c.c.. \quad (16)$$

Transformation \times Inverse transformation = **1**, as

$$\int_{-\infty}^{\infty} dk e^{ik(x-x')} = \delta(x - x'), \quad \int_{-\infty}^{\infty} d\omega e^{i\omega(t-t')} = \delta(t - t'), \quad (17)$$

δ - Dirac's delta.

Fourier-modes: $\hat{u}(k, \omega) e^{i(kx - \omega t)} \leftrightarrow$ monochromatic waves.

Amplitude: $|\hat{u}|$; Phase: $\Phi = kx - \omega t + \Phi_0$, $\hat{u} = |\hat{u}| e^{i\Phi}$.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Equations in sigma coordinates

Superposition principle

Method of Fourier \Leftrightarrow **superposition principle**, valid for **linear systems**.

$$u_t + cu_x = 0 \Rightarrow i(kc - \omega) \hat{u}(k, \omega), \hat{u}(k, \omega) \neq 0 \Rightarrow \quad (18)$$

$$\omega = ck, \text{ dispersion relation.} \quad (19)$$

General solution:

$$u(x, t) = \frac{1}{2\pi} \int dk e^{ik(x-ct)} \hat{u}(k) + c.c. \rightarrow \quad (20)$$

superposition (sum or integral) of elementary Fourier-modes.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Phase velocity

Speed of propagation of the phase of a monochromatic wave: **phase velocity**:

$$c_{ph} = \frac{\omega}{k}. \quad (21)$$

Dispersion: dependence $c = c(k) \Rightarrow$ simple wave is non-dispersive: $c_{ph} = c = \text{const.}$

Groupe velocity:

$$c_g = \frac{\partial \omega}{\partial k} \quad (22)$$

- speed of propagation of modulations = speed of propagation of information.

Simple wave: $c_{ph} = c_g$ (like acoustic or electromagnetic waves).

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Second-order wave equation

$$u_{tt} - c^2 u_{xx} = 0. \quad (23)$$

Same change of independent variables as in the 1st-order equation:

$$(x, t) \rightarrow (\xi_+, \xi_-) = (x + ct, x - ct)$$

$$u_{tt} - c^2 u_{xx} = 0 \rightarrow 4c^2 \frac{\partial^2 u}{\partial \xi_+ \partial \xi_-} = 0 \Rightarrow \quad (24)$$

General solution:

$$u = u_-(\xi_-) + u_+(\xi_+), \quad (25)$$

where $u_- + u_+$ - arbitrary functions, to be determined from initial conditions. (2nd order \Rightarrow 2 initial conditions required.)

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

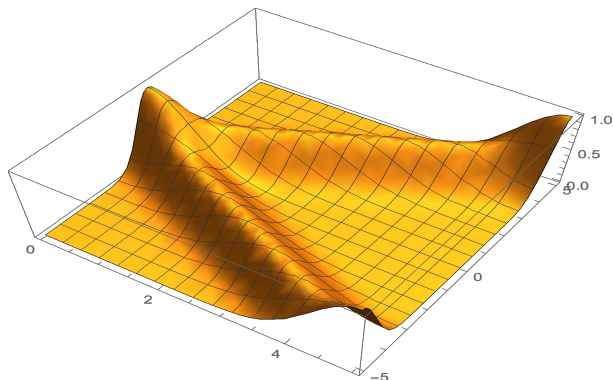
Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Spatio-temporal evolution of the initial localized perturbation



Solution in the domain $-5 < x < 5$, $0 < t < 5$. Initial Gaussian perturbation propagates along a **pair of characteristic** lines with slopes $\pm c$. Graphics by Mathematica[©]

Necessary mathematics

- Vector algebra
- Differential operations on scalar and vector fields
- Integration in 3D space
- Fourier analysis

Basic notions of wave dynamics

- Simple-wave equation**
- Dispersion, non-linearity

A crash course in fluid dynamics

- The perfect fluid
 - Governing equations
 - Euler - Lagrange duality
 - Energy and thermodynamics
 - Kelvin circulation theorem
- Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

- Primitive equations: Ocean
- Primitive equations: Atmosphere
- Equations in pressure coordinates
- Quasi-geostrophic approximation

Introducing the simplest dispersion

Dispersion - more derivatives.

In the case of **unidirectional propagation** - only odd-order derivatives to respect the symmetry of the initial equation with respect to reflexions. Simplest case: adding 3rd space derivative:

$$u_t + cu_x = 0 \rightarrow u_t + cu_x + \alpha u_{xxx} = 0 \quad \alpha = \text{const} \quad (26)$$

Corresponds to waves in shallow channels.

Dispersion relation:

$$\omega = ck - \alpha k^3$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

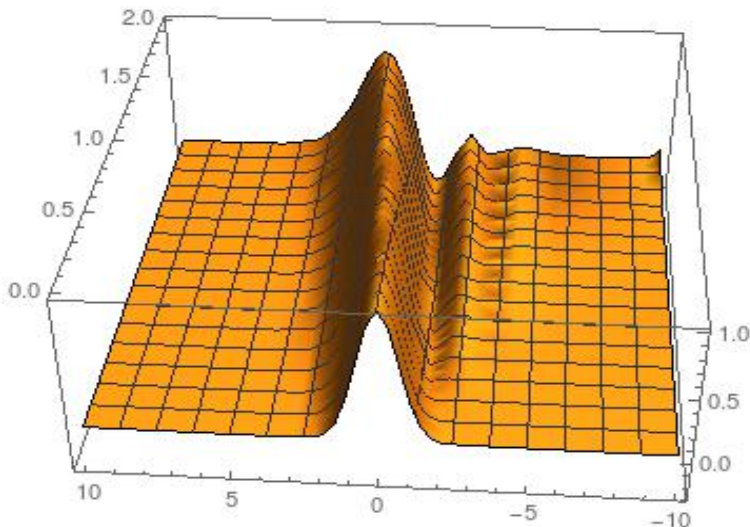
Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Spatio-temporal evolution of a localized initial perturbation, as follows from (26)



Necessary mathematics

- Vector algebra
- Differential operations on scalar and vector fields
- Integration in 3D space
- Fourier analysis

Basic notions of wave dynamics

- Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

- The perfect fluid
 - Governing equations
 - Euler - Lagrange duality
 - Energy and thermodynamics
 - Kelvin circulation theorem
- Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

- Primitive equations: Ocean
- Primitive equations: Atmosphere
 - Equations in pressure coordinates
 - Quasi-geostrophic approximation

Non-linearity

$$u_t + cu_x = 0 \rightarrow u_t + uu_x + cu_x = 0 \Rightarrow \quad (27)$$

no more superposition principle. Produces steepening and **wave breaking**.

Qualitative explanation : $c \rightarrow c + u \Rightarrow$ the larger the amplitude the larger the speed: a maximum moves faster than surrounding and “catches up” with the preceding part.

Korteweg - deVries equation: mutual compensation of dispersion and nonlinearity

Dispersion + non-linearity:

$$u_t + cu_x = 0 \rightarrow u_t + uu_x + cu_x + \alpha u_{xxx} = 0 \quad (28)$$

Produces **steady solitary waves**.

Necessary mathematics

- Vector algebra
- Differential operations on scalar and vector fields
- Integration in 3D space
- Fourier analysis

Basic notions of wave dynamics

- Simple-wave equation
- Dispersion, non-linearity

A crash course in fluid dynamics

- The perfect fluid
 - Governing equations
 - Euler - Lagrange duality
 - Energy and thermodynamics
 - Kelvin circulation theorem
- Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

- Primitive equations: Ocean
- Primitive equations: Atmosphere
- Equations in pressure coordinates
- Quasi-geostrophic approximation

Equations of motion

Eulerian description: in terms of fluid velocity field $\mathbf{v}(\mathbf{x}, t)$, and scalar density and pressure fields $\rho(\mathbf{x}, t)$, $P(\mathbf{x}, t)$, defined at each point \mathbf{x} of the volume occupied by the fluid at any time t .

Euler equations

Local conservation of momentum in the presence of forcing \mathbf{F} :

$$\rho \left(\frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v} \right) = -\nabla P + \mathbf{F}, \quad (29)$$

Continuity equation

Local conservation of mass:

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0. \quad (30)$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Vertical coordinate systems

Equation of state: baroclinic fluid

Fluid: **thermodynamical system** \Rightarrow equation of state relating P and ρ and closing the system (29), (30) (4 equations for 5 dependent variables).

General equation of state:

$$P = P(\rho, s), \quad (31)$$

$s(\mathbf{x}, t)$ is entropy per unit mass \Rightarrow evolution equation for s required. **Perfect fluid:**

$$\frac{\partial s}{\partial t} + \mathbf{v} \cdot \nabla s = 0. \quad (32)$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Equation of state: barotropic fluid

$$P = P(\rho) \leftrightarrow s = \text{const}, \quad (33)$$

sufficient to close the system (29), (30).

Particular case: **incompressible fluid**. Conservation of volume per unit mass \Rightarrow zero divergence:

$$\nabla \cdot \mathbf{v} = 0, \Rightarrow \quad (34)$$

$$\frac{\partial \rho}{\partial t} + \mathbf{v} \cdot \nabla \rho = 0, \quad \text{and} \quad \nabla \cdot (\mathbf{v} \cdot \nabla \mathbf{v}) = -\nabla \cdot \left(\frac{\nabla P}{\rho} \right) \Rightarrow \quad (35)$$

Pressure entirely determined by density and velocity distributions.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Lagrangian view of the fluid: momentum balance

Fluid \equiv ensemble of fluid parcels with time-dependent positions $\mathbf{X}(\mathbf{x}_0, t)$, $\mathbf{X}(\mathbf{x}_0, 0) = \mathbf{x}$.

Euler - Lagrange duality: continuity of the fluid \Rightarrow any point in the flow \mathbf{x} is, at the same time, a position of some fluid parcel \Rightarrow Eulerian velocity at the point $\mathbf{v}(\mathbf{x}) =$ velocity of the parcel $\mathbf{v}(\mathbf{X}, t) = \frac{d\mathbf{X}}{dt} \equiv \dot{\mathbf{X}}$. **Lagrangian (material) derivative** in Eulerian terms by chain differentiation:

$$\frac{d}{dt} = \frac{\partial}{\partial t} + \frac{\partial \mathbf{x}}{\partial t} \cdot \nabla \equiv \frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla. \quad (36)$$

\Rightarrow Newton's second law for the parcel

$$\rho(\mathbf{X}, t) \frac{d^2 \mathbf{X}}{dt^2} = -\nabla_{\mathbf{x}} P(\mathbf{X}, t) + \mathbf{F}, \quad (37)$$

\Leftrightarrow Euler equation (29).

Necessary mathematics

- Vector algebra
- Differential operations on scalar and vector fields
- Integration in 3D space
- Fourier analysis

Basic notions of wave dynamics

- Simple-wave equation
- Dispersion, non-linearity

A crash course in fluid dynamics

- The perfect fluid
 - Governing equations
 - Euler - Lagrange duality**
 - Energy and thermodynamics
 - Kelvin circulation theorem
- Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

- Primitive equations: Ocean
- Primitive equations: Atmosphere
 - Equations in pressure coordinates
 - Quasi-geostrophic approximation

Lagrangian view of the fluid: mass balance

Mass conservation in Lagrangian terms:

$$\rho_i(\mathbf{x})d^3\mathbf{x} = \rho(\mathbf{X}, t)d^3\mathbf{X}, \leftrightarrow \rho_i(\mathbf{x}) = \rho(\mathbf{X}, t)\mathcal{J} \quad (38)$$

where ρ_i is the initial distribution of density, and $d^3\mathbf{x}$ and $d^3\mathbf{X}$ are initial and current elementary volumes. The Jacobi determinant (Jacobian) in this formula is defined as the determinant:

$$\mathcal{J} = \begin{vmatrix} \frac{\partial X}{\partial x} & \frac{\partial X}{\partial y} & \frac{\partial X}{\partial z} \\ \frac{\partial Y}{\partial x} & \frac{\partial Y}{\partial y} & \frac{\partial Y}{\partial z} \\ \frac{\partial Z}{\partial x} & \frac{\partial Z}{\partial y} & \frac{\partial Z}{\partial z} \end{vmatrix} \equiv \frac{\partial(X, Y, Z)}{\partial(x, y, z)}$$

Incompressibility in Lagrangian terms: $\mathcal{J} = 1$. Taking Lagrangian time-derivative of this relation, we obtain the incompressibility condition of zero velocity divergence in Eulerian terms. Advection of entropy (32) \Leftrightarrow conservation of entropy by each fluid parcel $\dot{s} = 0$.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

1st principle of thermodynamics

Reversible processes in one-phase systems:

$$\delta\epsilon = T\delta s - P\delta v, \quad (39)$$

ϵ - internal energy per unit mass, $v = \frac{1}{\rho}$ - specific volume. Enthalpy per unit mass: $h = \epsilon + Pv \Rightarrow$

$$\delta h = T\delta s + v\delta P. \quad (40)$$

Energy density: sum of kinetic and internal parts:

$$e = \frac{\rho \mathbf{v}^2}{2} + \rho\epsilon. \quad (41)$$

Local conservation of energy :

$$\frac{\partial e}{\partial t} + \nabla \cdot \left[\rho \mathbf{v} \left(\frac{\mathbf{v}^2}{2} + h \right) \right] = 0. \quad (42)$$

Barotropic fluid:

$$\delta h = \frac{\delta P}{\rho} \Rightarrow \frac{\nabla P}{\rho} = \nabla h. \quad (43)$$

Necessary mathematics

Vector algebra

Differential operations on

scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations:

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Kelvin theorem

Circulation of velocity around a contour Γ consisting of fluid parcels, and moving with the fluid:

$$\gamma = \int_{\Gamma} \mathbf{v} \cdot d\mathbf{l} = \int_{S_{\Gamma}} (\nabla \wedge \mathbf{v}) \cdot d\mathbf{l}, \quad (44)$$

Kelvin theorem states that

- ▶ for barotropic fluids

$$\frac{d\gamma}{dt} = 0, \quad (45)$$

- ▶ for baroclinic fluids

$$\frac{d\gamma}{dt} = - \int_{\Gamma} \frac{\nabla P}{\rho} \cdot d\mathbf{l}. \quad (46)$$

Proof: direct calculation of the time-derivative of the circulation using the equations of motion, and the Lagrangian nature of Γ .

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Perfect vs real fluids

Perfect fluid approximation: **macroscopic** fluxes of mass, momentum and energy. Real fluids: corrections to these fluxes due to **molecular transport**. Simplest way to include them: **flux-gradient relations** following from **Le Chatelier principle**: molecular fluxes tend to restore the thermodynamical equilibrium. For any thermodynamical variable A

$$\mathbf{f}_A = -k_A \nabla A,$$

where \mathbf{f}_A is related molecular flux, and k_A is molecular transport coefficient.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Equations in pressure coordinates

Equations in pressure coordinates

Viscosity, diffusivity, and thermal conductivity

- ▶ Viscosity corrections to the Euler equation in the incompressible case, giving the **Navier - Stokes** equation

$$\frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v} = -\frac{\nabla P}{\rho} + \nu \nabla^2 \mathbf{v}, \quad \nabla \cdot \mathbf{v} = 0. \quad (47)$$

- ▶ Diffusivity corrections to the continuity equation

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = D \nabla^2 \rho. \quad (48)$$

- ▶ Thermal conductivity corrections to the heat/temperature advection giving the heat equation

$$\frac{\partial T}{\partial t} + \mathbf{v} \cdot \nabla T = \chi \nabla^2 T. \quad (49)$$

ν, D, χ are kinematic viscosity, diffusivity, and thermo-conductivity, the molecular transport coefficients for momentum, mass, and energy, respectively, all with dimension $\left[\frac{L^2}{T} \right]$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Dimensional/scale analysis. Reynolds number

Molecular transport coefficients: dimensional, value varies with changes if units. Only *non-dimensional parameters* are relevant. Typical space and velocity scales in the incompressible fluid flow: L , U . Time-scale $T = L/U$. Pressure scale: ρU^2 .

Scaled NS equation:

$$\frac{U^2}{L} \left(\frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v} + \nabla P \right) = \frac{U \nu}{L^2} \nabla^2 \mathbf{v} \rightarrow \quad (50)$$

Non-dimensional NS equation

$$\frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v} = -\nabla P + \frac{1}{Re} \nabla^2 \mathbf{v} \quad (51)$$

$Re = \frac{UL}{\nu}$ - **Reynolds number**, the true measure of viscosity. Similar, Pecklet number for diffusivity.

Necessary mathematics

- Vector algebra
- Differential operations on scalar and vector fields
- Integration in 3D space
- Fourier analysis

Basic notions of wave dynamics

- Simple-wave equation
- Dispersion, non-linearity

A crash course in fluid dynamics

- The perfect fluid
 - Governing equations
 - Euler - Lagrange duality
 - Energy and thermodynamics
 - Kelvin circulation theorem
- Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

- Primitive equations: Ocean
- Primitive equations: Atmosphere
 - Equations in pressure coordinates
 - Vertical coordinate transformation

Motion in a rotating frame

Material point in a frame rotating with angular velocity Ω :

$$m \frac{d\mathbf{v}}{dt} + 2m\Omega \wedge \mathbf{v} + m\Omega \wedge (\Omega \wedge \mathbf{x}) = \mathbf{F}, \quad \mathbf{v} = \frac{d\mathbf{x}}{dt} \quad (52)$$

m - mass, \mathbf{x} -current position of the point, \mathbf{F} - sum of forces acting on the point

Euler equations in the rotating frame +gravity:

Fluid under the influence of gravity: $m \rightarrow \rho$,
 $\frac{d}{dt} \rightarrow \frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla$, forces: pressure + gravity \Rightarrow

$$\frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v} + 2\Omega \wedge \mathbf{v} = -\frac{\nabla P}{\rho} + \mathbf{g}^* \quad (53)$$

Effective gravity: gravity + centrifugal acceleration (also potential)

$$\mathbf{g}^* = \mathbf{g} + \Omega \wedge (\Omega \wedge \mathbf{x}) \quad (54)$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

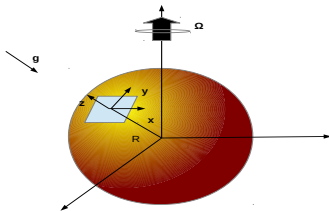
Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Geostrophic approximation

Tangent plane approximation



$$\frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v} + f \hat{\mathbf{z}} \wedge \mathbf{v} = -\frac{\nabla P}{\rho} + \mathbf{g} \quad (55)$$

f - plane: $f = \text{const}$; **β - plane:** $f = f + \beta y$; **f - Coriolis parameter:** $f = 2\Omega \sin \phi$, where ϕ - latitude

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Equations in pressure coordinates

Developing approximations

Hydrostatics. Stratification

The state of rest $\mathbf{v} \equiv 0$ is solution of (55) if **hydrostatic equilibrium** holds:

$$0 = -\frac{\nabla P}{\rho} + \mathbf{g}$$

The continuity equation:

$$\frac{d\rho}{dt} + \rho \nabla \cdot \mathbf{v} = 0$$

is satisfied by time-independent ρ in a state of rest.

Statically stable states: $\rho = \rho_0(z)$, $\rho'_0(z) \leq 0 \rightarrow$

$$P = P_0(z) = - \int dz g \rho_0(z)$$

Dependence of ρ_0 on z is called **stratification**. Surfaces of constant ρ : **isopycnals**.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

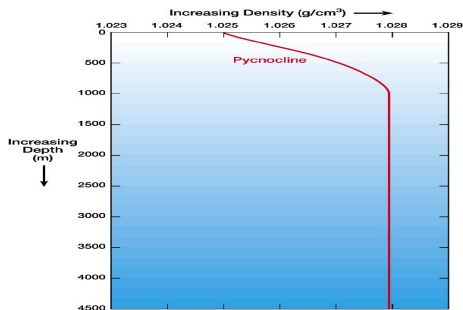
Primitive equations: Atmosphere

Equations in pressure coordinates

Geostrophic approximation

Oceanic stratification

Typical density profile:



$$\rho(\vec{x}, t) = \rho_0 + \rho_s(z) + \sigma(x, y, z; t), \quad \rho_0 \gg \rho_s \gg \sigma. \quad (56)$$

Hydrostatic approximation for large-scale motions:

$$g\rho + \partial_z P = 0, \Rightarrow P = P_0 + P_s(z) + \pi(x, y, z; t), \quad (57)$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Further approximations.

Boussinesq approximation

Deviations of density from ρ_0 neglected in the horizontal

→

$$\frac{\partial \mathbf{v}_h}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v}_h + f \hat{\mathbf{z}} \wedge \mathbf{v}_h = -\frac{\nabla_h \pi}{\rho} \approx -\nabla_h \phi, \quad (58)$$

where $\phi = \frac{\pi}{\rho_0}$ - **geopotential**.

Incompressibility of water

Continuity equation splits in two:

$$\nabla \cdot \mathbf{v} = 0, \quad \mathbf{v} = \mathbf{v}_h + \hat{\mathbf{z}} w. \quad (59)$$

$$\partial_t \rho + \mathbf{v} \cdot \nabla \rho = 0. \quad (60)$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Boussinesq approximation

Full set of oceanic PE

$$\frac{\partial \mathbf{v}_h}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v}_h + f \hat{\mathbf{z}} \wedge \mathbf{v}_h = -\frac{\nabla_h \pi}{\rho} \equiv -\nabla_h \phi, \quad (61)$$

$$\partial_t \sigma + \mathbf{v} \cdot \nabla \sigma + w \rho'_s(z) = 0. \quad (62)$$

$$g \frac{\sigma}{\rho_0} = -\partial_z \phi, \quad \nabla_h \cdot \mathbf{v}_h + \partial_z w = 0, \quad (63)$$

Remark

Hydrostatic approximation \leftrightarrow anisotropic scaling proper for mesoscale motions:

$$W \ll U, \quad H \ll L, \quad \frac{W}{H} \sim \frac{U}{L}$$

where L , H and U , W are horizontal and vertical spatial and velocity scales, respectively.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Vertical boundary conditions

Most often sufficient for our purposes: **rigid lid and flat bottom**:

$$w|_{z=0} = w|_{z=H} = 0 \quad (64)$$

Non-trivial bathymetry : fluid parcels follow the bottom profile

$$w|_{z=b(x,y)} = \frac{db}{dt} = \mathbf{v} \cdot \nabla b$$

Free surface: fluid parcels move with the surface:

$$w|_{z=h(x,y;t)} = \frac{dh}{dt} = \frac{\partial h}{\partial t} + \mathbf{v} \cdot \nabla h$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

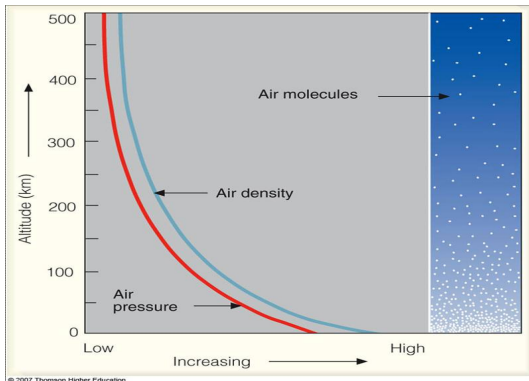
Primitive equations:

Atmosphere

Equations in pressure coordinates

Quasi-geostrophic approximation

Atmosphere: pressure coordinates



Altitude \leftrightarrow Pressure \Rightarrow vertical coordinate.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Equations in pressure coordinates

Thermodynamics of the dry atmosphere

Equation of state - ideal gas:

$$P = \rho RT, \quad c_{P,V} = T \left(\frac{\partial s}{\partial T} \right)_{P,V} = \text{const}, \quad c_p - c_v = R. \quad (65)$$

Entropy:

$$s = c_p \ln T - R \ln P + \text{const}. \quad (66)$$

Adiabatic process:

$$s = \text{const} \Rightarrow c_p \frac{dT}{T} - R \frac{dP}{P} = 0, \Rightarrow T = T_s \left(\frac{P}{P_s} \right)^{\frac{R}{c_p}}. \quad (67)$$

Potential temperature :

$$\theta = T \left(\frac{P_s}{P} \right)^{\frac{R}{c_p}}, \quad s = c_p \ln \theta + \text{const}. \quad (68)$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Equations in pressure coordinates

Geopotential and hydrostatics

Geopotential variation: work to lift a unit mass against gravity: $\delta\phi = g\delta z$.

$z = z(p)$ becomes a **thermodynamical variable**.

Hydrostatic approximation:

$$\delta\phi = -\frac{RT}{P}\delta P \Rightarrow \quad (69)$$

$$\frac{\partial\phi}{\partial P} = -\frac{RT}{P} = -\frac{1}{\rho}. \quad (70)$$

Useful relation for **small variations** ρ , P , θ with respect to background ρ_0 , P_0 , θ_0 :

$$\theta = \theta_0 \left[\frac{\left(1 - \frac{R}{c_p}\right) P}{P_0} - \frac{\rho}{\rho_0} \right] \quad (71)$$

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Equations in isentropic coordinates

Elimination of ρ in Euler equations

"Triangular" relation :

$$\left(\frac{\partial P}{\partial x}\right)_z \left(\frac{\partial x}{\partial z}\right)_P \left(\frac{\partial z}{\partial P}\right)_x = -1 \Rightarrow \quad (72)$$

$$\left(\frac{\partial P}{\partial x}\right)_z = - \left(\frac{\partial P}{\partial z}\right)_x \left(\frac{\partial z}{\partial x}\right)_P = \rho \left(\frac{\partial \phi}{\partial x}\right)_P. \quad (73)$$

Incompressibility in pressure coordinates

Lagrangian volume element in pressure coordinates:

$$\rho dx dy dz = -\frac{1}{g} dx dy dP \quad (74)$$

Mass conservation \Rightarrow Volume conservation in P .

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Equations in pressure coordinates

Adiabatic primitive equations

Equations of motion

$$\operatorname{div}(\mathbf{v}) = \nabla_h \cdot \mathbf{v}_h + \partial_p \omega = 0, \quad \omega = \frac{dP}{dt}. \quad (75)$$

$$\frac{\partial \mathbf{v}_h}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v}_h + f \hat{\mathbf{z}} \wedge \mathbf{v}_h = -\nabla_h \phi. \quad (76)$$

$$\partial_t \theta + \mathbf{v} \cdot \nabla \theta = 0. \quad (77)$$

$$\frac{\partial \phi}{\partial P} = -\frac{RT}{P} = -\frac{R}{P} \left(\frac{P}{P_s} \right)^{\frac{R}{c_p}} \theta. \quad (78)$$

Boundary conditions

Bottom: ground \equiv **free surface** in terms of pressure, geopotential fixed.

Top: rigid lid \equiv fixed value of pressure, e.g. tropopause.

Necessary mathematics

Vector algebra

Differential operations on scalar and vector fields

Integration in 3D space

Fourier analysis

Basic notions of wave dynamics

Simple-wave equation

Dispersion, non-linearity

A crash course in fluid dynamics

The perfect fluid

Governing equations

Euler - Lagrange duality

Energy and thermodynamics

Kelvin circulation theorem

Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

Primitive equations: Ocean

Primitive equations: Atmosphere

Atmosphere

Equations in pressure coordinates

Vertical coordinate systems

Boussinesq approximation for atmosphere

Varying background density in atmosphere: $\rho_0 = \rho_0(z)$.

Boussinesq approximation in x, y, z coordinates, with $\rho = \rho_0(z) + \tilde{\rho}$, $P = P_0(z) + \tilde{p}$, $\theta = \theta_0(z) + \tilde{\theta}$, ($\tilde{\dots}$) omitted below:

$$\frac{\partial \mathbf{v}_h}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v}_h + f \hat{\mathbf{z}} \wedge \mathbf{v}_h = -\nabla_h \phi, \quad (79)$$

with geopotential $\phi = \frac{p}{\rho_0}$. Hydrostatics:

$$-\frac{\partial \phi}{\partial z} - \frac{p}{\rho_0^2} \frac{\partial \rho_0}{\partial z} - g \frac{\rho}{\rho_0} = 0. \quad (80)$$

Equation of state (ideal gas) + (71) \rightarrow

$$-\frac{\partial \phi}{\partial z} + b = 0, \quad (81)$$

$b = g \frac{\theta}{\theta_0}$ - **buoyancy**, $\frac{\partial b}{\partial t} + \mathbf{v} \cdot \nabla b = 0$ for adiabatic motions.

Continuity equation \rightarrow **anelastic equation**:

$$\nabla \cdot (\rho_0(z) \mathbf{v}) = 0.$$

Necessary mathematics

- Vector algebra
- Differential operations on scalar and vector fields
- Integration in 3D space
- Fourier analysis

Basic notions of wave dynamics

- Simple-wave equation
- Dispersion, non-linearity

A crash course in fluid dynamics

- The perfect fluid
 - Governing equations
 - Euler - Lagrange duality
 - Energy and thermodynamics
 - Kelvin circulation theorem
- Real fluids: incorporating molecular transport

Hydrodynamics on a tangent plane to a rotating planet

Primitive equations for the ocean and the atmosphere

- Primitive equations: Ocean
- Primitive equations: Atmosphere
- Equations in pressure coordinates
- Boussinesq approximation