1. Modeling large-scale atmospheric and oceanic motions: Primitive equations in Geophysical Fluid Dynamics

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Mathematics of the atmosphere and oceans, SUSTECH, 2023 Mathematics of the atmosphere and oceans 1

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Primitive equations (PE)

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First conclusions.

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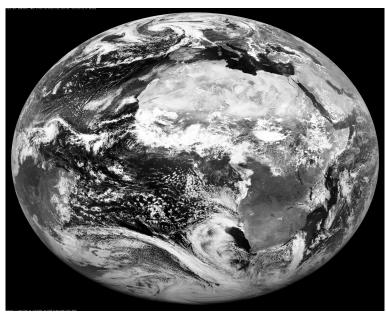
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Geophysical Fluid Dynamics : space view



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GFD : what's that?

Hydrodynamics in all its complexity plus :

- Rotating frame
- Variable temperature/density effects
- Spherical geometry (large- and meso-scales)
- Fluid in the complex domains (coasts, topography/bathymetry)
- Multi-phase fluid (water vapor, ice)

But !

Some of these additional effects, like fast rotation, allow to simplify the analysis

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Scales :

Horizontal scales

- Large : planetary 10⁴ km
- Medium : atmosphere synoptic, 10³ km; ocean meso-scale 10 - 10² km
- Small : atmosphere meso-scale 1 − 10 km; ocean sub-meso scale 1 km
- Very small : meters

Scales of interest : large and medium.

Vertical scales

Synoptic motions in the atmosphere : whole troposphere, or its significant part, O(10 km).

Meso-scale motions in the ocean : whole oceanic depth, or its significant part, $\mathcal{O}(1km)$.

 \Rightarrow Strong disparity between horizontal and vertical scales.

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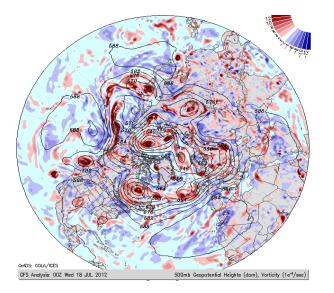
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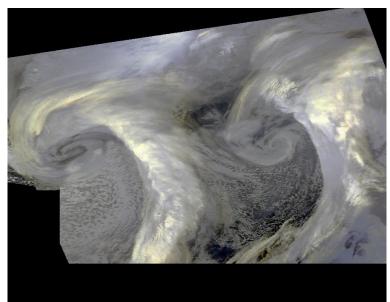
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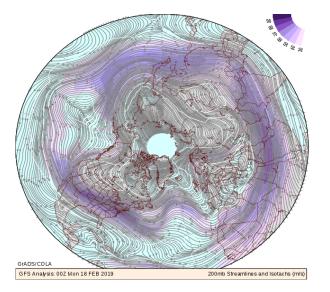
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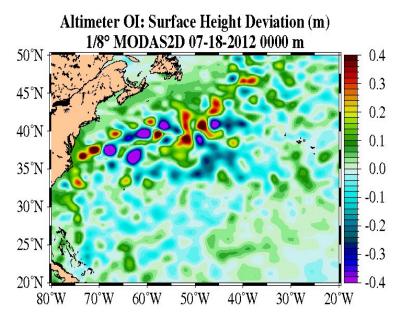
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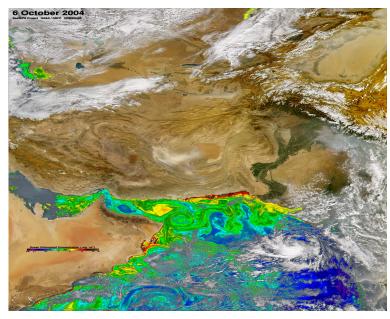
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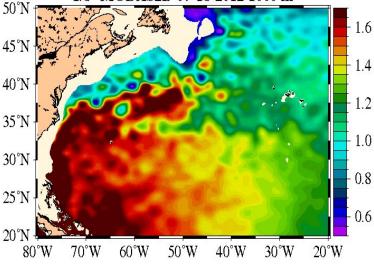
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Oceanic jet in satellite data

Altimeter OI: Steric Height Anomaly (m) 1/8° MODAS2D 07-18-2012 1000 m



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Where the governing equations come from :

- ► Mechanical system ⇒ Newton's 2nd law ↔ momentum conservation.
- ► Continuous medium ⇒ local mass conservation
- ► Thermodynamical system ⇒ 1st and 2nd laws of thermodynamics, equation of state
- Dissipative effects \Rightarrow flux-gradient relations.

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Dynamics of the perfect fluid : Lagrange's view

Description in terms of positions of fluid parcels X(a, t), along their trajectories, where a = (a, b, c) are initial positions (Lagrangian labels) and $v(X, t) = \frac{dX}{dt} \equiv \dot{X}$ is their velocity.

Newton's 2nd law :

$$\rho(\boldsymbol{X},t)\frac{d^{2}\boldsymbol{X}}{dt^{2}} = -\frac{\partial P(\boldsymbol{X},t)}{\partial \boldsymbol{X}} \equiv -\nabla_{\boldsymbol{X}}P(\boldsymbol{X},t), \qquad (1)$$

where ρ and P are density and pressure in the fluid. Mass conservation :

$$\rho_i(\boldsymbol{a})d^3\boldsymbol{a} = \rho(\boldsymbol{X},t)d^3\boldsymbol{X}, \leftrightarrow \rho_i(x) = \rho(\boldsymbol{X},t)\mathcal{J}(\boldsymbol{X},\boldsymbol{a}) \quad (2)$$

where ρ_i is initial distribution of density, $\mathcal{J}(\mathbf{X}, \mathbf{a}) = \frac{\partial(\mathbf{X}, \mathbf{Y}, \mathbf{Z})}{\partial(\mathbf{a}, b, c)}$ is the Jacobi determinant (Jacobian). $\frac{d\mathcal{J}}{dt} = \frac{\rho_i}{\rho} \nabla_{\mathbf{X}} \cdot \dot{\mathbf{X}}$ No heat exchange between parcels :

$$\frac{ds}{dt} = 0, \tag{3}$$

where $s(\mathbf{X}, t)$ is the entropy of the parcel.

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Dynamics of the perfect fluid : Euler's view

Description in terms of velocity, density and pressure fields at a fixed point of space : $\mathbf{v}(\mathbf{x}, t)$, $\rho(\mathbf{x}, t)$, $P(\mathbf{x}, t)$. Euler-Lagrange duality : $\forall \mathbf{x}, t \exists \mathbf{X}(t) = \mathbf{x}$. Chain differentiation \rightarrow

$$\frac{d}{dt} = \frac{\partial}{\partial t} + \mathbf{v} \cdot \boldsymbol{\nabla}. \tag{4}$$

Newton's 2nd law :

$$\rho\left(\frac{\partial \boldsymbol{v}}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla} \boldsymbol{v}\right) = -\boldsymbol{\nabla} \boldsymbol{P}.$$

Continuity (mass conservation) equation :

$$\frac{\partial \rho}{\partial t} + \boldsymbol{\nabla} \cdot (\rho \boldsymbol{v}) = 0 \leftrightarrow \frac{d}{dt} (\rho \mathcal{J}) = \frac{d\rho}{dt} \mathcal{J} + \rho \frac{d\mathcal{J}}{dt} = \frac{d\rho_i}{dt} = 0$$
(6)

Momentum conservation :

$$\frac{\partial \rho \mathbf{v}}{\partial t} + \boldsymbol{\nabla} \cdot (\rho \mathbf{v} \otimes \mathbf{v} + P \mathbb{1}) = 0$$
 (7)

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Adding thermodynamics : equation of state General equation of state

$$P = P(\rho, s),$$

where s - entropy per unit mass;

Barotropic (isentropic) fluid :

$$s = {
m const} \ \Rightarrow P = P(
ho),$$

Baroclinic fluid :

$$P = P(\rho, s), \tag{10}$$

Incompressible - particular case of barotropic- fluid :

$$\mathcal{J} = 1 \leftrightarrow \boldsymbol{\nabla} \cdot \boldsymbol{v} = 0 \Rightarrow \frac{d\rho}{dt} = \frac{\partial\rho}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla}\rho = 0. \quad (11)$$

 \Rightarrow pressure is not independent variable.

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Thermodynamics : reminder 1st principle, "dry" thermodynamics

$$d\epsilon = Tds - Pdv, \tag{12}$$

where ϵ - internal energy and $v = \frac{1}{\rho}$ - volume per unit mass. Enthalpy per unit mass : $h = \epsilon + Pv$:

$$dh = Tds + vdP.$$

Energy density of the fluid :

$$e = \frac{\rho \mathbf{v}^2}{2} + \rho \epsilon.$$

Local conservation of energy :

$$\frac{\partial \boldsymbol{e}}{\partial t} + \boldsymbol{\nabla} \cdot \left[\rho \boldsymbol{v} \left(\frac{\boldsymbol{v}^2}{2} + h \right) \right] = 0.$$
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Volume-preserving diffeomorphisms

Flow of the incompressible fluid in the domain \mathcal{D} in the absence of singularities \rightarrow bijective mapping $\boldsymbol{a} \mapsto \boldsymbol{X}(\boldsymbol{a}, t)$, $a, X \in \mathcal{D}$ of \mathcal{D} onto itself : volume-preserving diffeomorphism belonging to infinite-dimensional Lie group. Lagrangian equations of motion : Euler-Lagrange equations

$$\frac{d}{dt}\frac{\delta \mathcal{L}}{\delta \dot{\boldsymbol{X}}} - \frac{\delta \mathcal{L}}{\delta \boldsymbol{X}} = 0$$

for variational (Hamilton's) principle with action

$$S = \int dt \int d^{3}\mathbf{a}\mathcal{L}; \ \mathcal{L} = \rho_{i}(\mathbf{a})\frac{\dot{\mathbf{X}}^{2}(\mathbf{a},t)}{2} + P(\mathbf{a},t)\left(\mathcal{J}(\mathbf{X},\mathbf{a}) - 1\right)$$
(18)

geodesic equations on the group manifold, after relabeling $\boldsymbol{a} \rightarrow \boldsymbol{\tilde{a}}$: $\rho_i(\boldsymbol{a}) d^3 \boldsymbol{a} = d^3 \boldsymbol{\tilde{a}}$ (mass-weighted labels). *P* -Lagrange multiplier to the incompressibility constraint.

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Dissipative phenomena as molecular fluxes

Perfect fluid : local conservation laws with macroscopic fluxes of related quantities. Dissipation : correction of the macroscopic fluxes of

- momentum
- mass
- internal energy (heat)

by corresponding molecular fluxes, calculated from the flux - gradient relations (LeChatelier principle) :

$$\boldsymbol{f}_{\boldsymbol{A}} = -k_{\boldsymbol{A}}\boldsymbol{\nabla}\boldsymbol{A}, \tag{19}$$

A - a thermodynamical variable, f_A - corresponding molecular flux, k_A - transport coefficient.

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Introducing viscosity

Stress tensor in momentum conservation corrected by viscous stresses - (density of) molecular momentum flux :

$$ho oldsymbol{v} \otimes oldsymbol{v} o
ho oldsymbol{v} \otimes oldsymbol{v} - \mu oldsymbol{
abla} \otimes oldsymbol{v}$$

 $\mu = \rho \nu$, μ - dynamic, ν - kinematic viscosities Incompressible case : Navier -Stokes (NS) equation

$$\frac{\partial \boldsymbol{v}}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla} \boldsymbol{v} = -\frac{\boldsymbol{\nabla} P}{\rho} + \nu \boldsymbol{\nabla}^2 \boldsymbol{v}, \ \boldsymbol{\nabla} \cdot \boldsymbol{v} = 0.$$
(21)

Reynolds' number

Dimensionless form of the NS equation :

$$\frac{\partial \boldsymbol{v}}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla} \boldsymbol{v} = -\frac{\boldsymbol{\nabla} P}{\rho} + \frac{1}{Re} \boldsymbol{\nabla}^2 \boldsymbol{v}, \qquad (22)$$

where $Re = UL/\nu$, U, L -typical scales of velocity and length. Synoptic motions far from the boundaries : $Re \rightarrow \infty$. Mathematics of the atmosphere and oceans 1

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Diffusivity, thermal conductivity

Molecular fluxes of mass and heat :

$$-D\boldsymbol{\nabla}\rho, \quad -\kappa\boldsymbol{\nabla}T$$

Corrected continuity equation :

$$\frac{\partial \rho}{\partial t} + \boldsymbol{\nabla} \cdot (\rho \boldsymbol{\nu}) = D \boldsymbol{\nabla}^2 \rho.$$

Equation of heat/temperature

$$\frac{\partial T}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla} T = \chi \boldsymbol{\nabla}^2 T.$$

Non-dimensional form \rightarrow Péclet numbers Pe = UL/D, $Pe = UL/\chi$, small for synoptic motions far from the boundaries.

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Motion in the rotating frame

Material point in the rotating frame :

$$mrac{doldsymbol{v}}{dt}+2moldsymbol{\Omega}\wedgeoldsymbol{v}+moldsymbol{\Omega}\wedge(oldsymbol{\Omega}\wedgeoldsymbol{x})=oldsymbol{F},$$

where $\boldsymbol{v} = \frac{d\boldsymbol{x}}{dt}$, and \boldsymbol{F} is a sum of external forces.

Euler equations with rotation Replacements $m \to \rho, \frac{d}{dt} \to \frac{\partial}{\partial t} + \mathbf{v} \cdot \nabla, \ \mathbf{F} \to -\nabla P$:

$$\frac{\partial \boldsymbol{v}}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla} \boldsymbol{v} + 2\boldsymbol{\Omega} \wedge \boldsymbol{v} + \boldsymbol{\Omega} \wedge (\boldsymbol{\Omega} \wedge \boldsymbol{x}) = -\frac{\boldsymbol{\nabla} P}{\rho}, \quad (27)$$

Including rotation in the variational principle

$$\mathcal{L} \rightarrow \mathcal{L} + \mathcal{R} \cdot \dot{\boldsymbol{X}}; \quad \mathcal{R} : \boldsymbol{\nabla}_{\boldsymbol{X}} \wedge \mathcal{R} = 2\boldsymbol{\Omega}.$$
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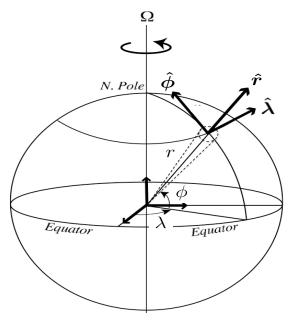
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Spherical coordinates



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Euler and continuity equations

$$\frac{dv_r}{dt} - \frac{v_\lambda^2 + v_\phi^2}{r} - 2\Omega \cos \phi v_\lambda + g = -\frac{1}{\rho} \partial_r P,$$

$$\frac{dv_\lambda}{dt} + \frac{v_r v_\lambda - v_\phi v_\lambda \tan \phi}{r} + 2\Omega \left(-\sin \phi v_\phi + \cos \phi v_r \right)$$

$$= -\frac{1}{\rho r \cos \phi} \partial_\lambda P,$$

$$\frac{dv_\phi}{dt} + \frac{v_r v_\phi + v_\lambda^2 \tan \phi}{r} + 2\Omega \sin \phi v_\lambda = -\frac{1}{\rho r} \partial_\phi P,$$

$$\frac{d\rho}{dt} + \rho \left[\frac{1}{r^2} \frac{\partial (r^2 v_r)}{\partial r} + \frac{1}{r \cos \phi} \left(\frac{\partial (\cos \phi v_\phi)}{\partial \phi} + \frac{\partial v_\lambda}{\partial \lambda} \right) \right],$$

$$\frac{d}{dt} = \frac{\partial}{\partial t} + v_r \partial_r + \frac{v_\phi}{r} \partial_\phi + \frac{v_\lambda}{r \cos \phi} \partial_\phi$$
Traditional approximation : green + red \rightarrow out,
 $r \rightarrow R = \text{const. centrifugal acceleration neglected}$

Non-traditional approximation : green \rightarrow out.

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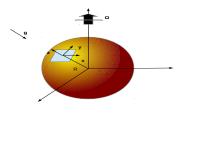
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Tangent plane approximation (traditional)



$$\frac{\partial \boldsymbol{v}}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla} \boldsymbol{v} + f \hat{\boldsymbol{z}} \wedge \boldsymbol{v} = -\frac{\boldsymbol{\nabla} P}{\rho} + \boldsymbol{g}^*$$
(29)

Coriolis parameter $f = 2\Omega \sin \phi$, \hat{z} - unit z, effective gravity :

$$\boldsymbol{g}^* = \boldsymbol{g} + \boldsymbol{\Omega} \wedge (\boldsymbol{\Omega} \wedge \boldsymbol{x})$$
 (30)

"Traditional" : $g^* \approx g$ (correction several %). f - plane : f = const; β - plane : $f = f + \beta y$; Mathematics of the atmosphere and oceans 1

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Hydrostatics. Stratification

State of rest $\mathbf{v} = (u, v, w) \equiv 0$ solution of (29) if hydrostatic equilibrium holds :

$$\mathsf{0}=-rac{oldsymbol{
abla}P}{
ho}+oldsymbol{g}$$

Continuity equation : satisfied by time-independent ρ in a state of rest.

$$\frac{d\rho}{dt} + \rho \boldsymbol{\nabla} \cdot \boldsymbol{v} = \boldsymbol{0}$$

Statically stable states : $\rho = \rho_0(z)$, $\rho'_0(z) \le 0$. Dependence of ρ_0 on z : stratification. Hydrostatic approximation : vertical accelerations small - valid for large scales \Leftrightarrow large masses.

$$\frac{dw}{dt} \ll g.$$

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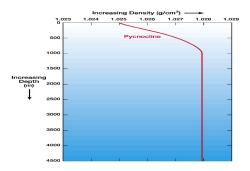
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Ocean : observations and approximations Typical averaged density profile :



 $\rho(\mathbf{x},t) = \rho_0 + \rho_s(z) + \sigma(x,y,z;t), \ \rho_0 \gg ||\rho_s|| \gg ||\sigma||. \ (32)$

Temperature (T) and salinity (S) : density variations $\delta \rho \propto (\delta T, \delta S)$. Hydrostatics :

$$P = P_0 + P_s(z) + \pi(x, y, z; t),$$
(33)

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Equations of motion Incompressibility and hydrostatics :

$$oldsymbol{
abla}\cdotoldsymbol{
u}=0,\quad g
ho+\partial_zP=0,$$

Euler equations in Boussinesq approximation :

$$\frac{\partial \boldsymbol{v}_h}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla} \boldsymbol{v}_h + f \hat{\boldsymbol{z}} \wedge \boldsymbol{v}_h = -\frac{\boldsymbol{\nabla}_h \pi}{\rho} \approx -\boldsymbol{\nabla}_h \phi.$$
(35)

 $\phi = \frac{\pi}{\rho_0}$ - geopotential, $\mathbf{v} = \mathbf{v}_h + \hat{\mathbf{z}}w \equiv u\hat{\mathbf{x}} + v\hat{\mathbf{y}} + w\hat{\mathbf{z}}$. Mass conservation :

$$\frac{d\rho}{dt} \equiv \partial_t \rho + \mathbf{v} \cdot \nabla \rho = 0. \tag{36}$$

Boundary conditions :

Rigid lid and flat bottom : $w|_{z=0} = w|_{z=H} = 0$. Non-trivial bathymetry b(x, y) and/or free-surface : h(x, y, t) :

$$w|_{z=b} = \frac{db}{dt}, \quad w|_{z=h} = \frac{dh}{dt}$$

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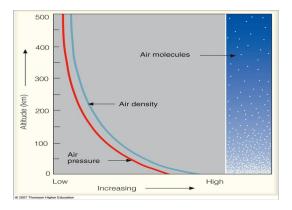
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Atmosphere : pressure coordinatees



One-to-one correspondence between altitude and pressure \rightarrow pressure as vertical coordinate.

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First conclusions.

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Thermodynamics of the dry atmosphere Equation of state - ideal gas :

$$P = \rho RT, \ c_{p,v} = T\left(\frac{\partial s}{\partial T}\right)_{P,V} = const, \ c_p - c_v = R.$$
(37)

Entropy :

$$s = c_p \ln T - R \ln P + const.$$

Adiabatic process :

$$s = \text{const} \Rightarrow c_p \frac{dT}{T} - R \frac{dP}{P} = 0, \Rightarrow T = T_s \left(\frac{P}{P_s}\right)^{\frac{R}{c_p}}.$$
 (39)

Potential temperature :

$$\theta = T\left(\frac{P_s}{P}\right)^{\frac{R}{c_p}}, s = c_p \ln \theta + \text{const.}$$
(40)

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Geopotential

 $d\phi = gdz$, where z = z(P) via hydrostatics, and thus becomes a thermodynamical variable.

Hydrostatics

$$\begin{split} d\phi &= -\frac{RT}{P}dP \rightarrow \\ \frac{\partial \phi}{\partial P} &= -\frac{RT}{P} = -\frac{1}{\rho} \rightarrow \\ \frac{\partial \phi}{\partial P} &= -\frac{R}{P}\left(\frac{P}{P_s}\right)^{\frac{R}{c_p}} \theta \end{split}$$

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Elimination of ρ in Euler equations "Triangular" relation :

$$\left(\frac{\partial P}{\partial x}\right)_{z}\left(\frac{\partial x}{\partial z}\right)_{P}\left(\frac{\partial z}{\partial P}\right)_{x} = -1 \Rightarrow$$

$$\left(\frac{\partial P}{\partial x}\right)_{z} = -\left(\frac{\partial P}{\partial z}\right)_{x} \left(\frac{\partial z}{\partial x}\right)_{P} = \rho\left(\frac{\partial \phi}{\partial x}\right)_{P}.$$
 (44)

Incompressibility in pressure coordinates

Lagrangian view : mass element \rightarrow volume element in pressure coordinates in hydrostatic approximation :

$$\rho dx dy dz = -\frac{1}{g} dx dy dP \tag{45}$$

Mass conservation \Rightarrow volume conservation with pressure as vertical coordinate \Rightarrow divergence-less velocity.

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Adiabatic equations of motion

$$div(\mathbf{v}) = \mathbf{\nabla}_h \cdot \mathbf{v}_h + \partial_p \omega = 0, \ \ \omega \equiv rac{dP}{dt}.$$

$$\frac{\partial \boldsymbol{v}_h}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla} \boldsymbol{v}_h + f \hat{\boldsymbol{z}} \wedge \boldsymbol{v}_h = -\boldsymbol{\nabla}_h \phi.$$
(47)

$$\partial_t \theta + \mathbf{v} \cdot \nabla \theta = 0. \tag{48}$$

$$\frac{\partial \phi}{\partial P} = -\frac{RT}{P} = -\frac{R}{P} \left(\frac{P}{P_s}\right)^{\frac{R}{c_p}} \theta.$$
(49)

Remark : Coordinates (x, y, P) curvilinear, but curvature terms are neglected in divergence and material derivative \leftrightarrow weak curvature of the isobars, mostly valid for synoptic motions.

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"Pseudo-height" coordinate

New vertical coordinate :

$$\bar{z} = z_0 \left(1 - \left(\frac{P}{P_s}\right)^{\frac{R}{c_p}} \right) \equiv z_0 \left(1 - \left(\frac{P}{P_s}\right)^{\frac{\gamma-1}{\gamma}} \right), \quad (50)$$

$$z_0 = rac{\gamma}{\gamma-1} rac{P_s}{g
ho_s} pprox 28 {
m km}.$$

Pseudo- density :

$$r: rd\bar{z} =
ho dz = -\frac{1}{g}dP.$$

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Mass conservation :

$$dxdydP = -gr(\bar{z})dxdyd\bar{z} \Rightarrow$$
 (53)

$$r\left(\boldsymbol{\nabla}_{h}\cdot\boldsymbol{v}_{h}+\frac{\partial\bar{w}}{\partial\bar{z}}\right)+\bar{w}\frac{\partial r}{\partial\bar{z}}=0,\quad\boldsymbol{v}=(\boldsymbol{v}_{h},\bar{w}=\frac{d\bar{z}}{dt}).$$
 (54)

Approximation $\bar{z} \ll z_0$:

$$\boldsymbol{\nabla}_{h} \cdot \boldsymbol{v}_{h} + \frac{\partial \bar{w}}{\partial \bar{z}} = -\frac{\bar{w}}{r} \frac{\partial r}{\partial \bar{z}} = \frac{\bar{w}}{(\gamma - 1)z_{0} \left(1 - \frac{\bar{z}}{z_{0}}\right)} \approx 0.$$
(55)

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Equations of motion

$$\frac{\partial \boldsymbol{v}_{h}}{\partial t} + \boldsymbol{v} \cdot \boldsymbol{\nabla} \boldsymbol{v}_{h} + f \hat{\boldsymbol{z}} \wedge \boldsymbol{v}_{h} = -\boldsymbol{\nabla}_{h} \phi, \qquad (56)$$
$$-g \frac{\theta}{\theta_{0}} + \frac{\partial \phi}{\partial \bar{\boldsymbol{z}}} = 0, \qquad (57)$$
$$\frac{\partial \theta}{\partial \theta} = -\boldsymbol{v}_{h} \cdot \boldsymbol{v}_{h} = -\boldsymbol{\nabla}_{h} \phi, \qquad (56)$$

$$\frac{\partial \theta}{\partial t} + \mathbf{v} \cdot \nabla \theta = 0; \quad \nabla \cdot \mathbf{v} = 0.$$
 (58)

Identical to oceanic PE with the exchange $\sigma \rightarrow -\theta$.

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Conservation of PV in the PE model

Absolute vorticity :

$$\boldsymbol{\zeta}_{a} = \boldsymbol{\zeta} + \hat{\mathbf{z}}f, \, \boldsymbol{\zeta} = \boldsymbol{\nabla} \wedge \boldsymbol{v}, \quad \boldsymbol{\nabla} \cdot \boldsymbol{\zeta}_{a} = \mathbf{0},$$
 (59)

Application of $\nabla \wedge$ to PE + "hydrodynamic identity" :

$$\boldsymbol{\nu} \cdot \boldsymbol{\nabla} \boldsymbol{\nu} = \frac{1}{2} \boldsymbol{\nabla} \boldsymbol{\nu}^2 - \boldsymbol{\nu} \wedge (\boldsymbol{\nabla} \wedge \boldsymbol{\nu})$$
(60)

ightarrow equation for ${oldsymbol{\zeta}}_a$:

$$\frac{d\boldsymbol{\zeta}_{a}}{dt} = \boldsymbol{\zeta}_{a} \cdot \boldsymbol{\nabla} \boldsymbol{\nu} + \frac{g}{\rho_{0}} \hat{\boldsymbol{z}} \wedge \boldsymbol{\nabla} \rho.$$
 (61)

Baroclinic torque

Last term in the r.h.s. : creation of horizontal vorticity by density gradients.

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Conservation of potential vorticity (PV)

$$PV \equiv q = \zeta_a \cdot \nabla \rho, \quad \frac{dq}{dt} = 0.$$
 (62)

PV anomaly (zero at the rest state) : $PVA = q - f \rho'_s(z)$. Proof by direct calculation :

$$\begin{aligned} \partial_t \left(\boldsymbol{\zeta}_a \cdot \boldsymbol{\nabla} \rho \right) &= \left(\partial_t \boldsymbol{\zeta}_a \right) \cdot \boldsymbol{\nabla} \rho + \boldsymbol{\zeta}_a \cdot \boldsymbol{\nabla} (\partial_t \rho) \\ &= \boldsymbol{\nabla} \rho \cdot \left(\boldsymbol{\nabla} \wedge \left(\boldsymbol{v} \wedge \boldsymbol{\zeta}_a \right) \right) - \boldsymbol{\zeta}_a \cdot \boldsymbol{\nabla} \left(\boldsymbol{v} \cdot \boldsymbol{\nabla} \rho \right) \\ &= -\boldsymbol{\nabla} \cdot \left(\boldsymbol{\nabla} \rho \wedge \left(\boldsymbol{v} \wedge \boldsymbol{\zeta}_a \right) \right) - \boldsymbol{\zeta}_a \cdot \boldsymbol{\nabla} \left(\boldsymbol{v} \cdot \boldsymbol{\nabla} \rho \right) \\ &= -\boldsymbol{\nabla} \cdot \left(\boldsymbol{v} \left(\boldsymbol{\zeta}_a \cdot \boldsymbol{\nabla} \rho \right) \right) + \boldsymbol{\nabla} \cdot \left(\boldsymbol{\zeta}_a \left(\boldsymbol{v} \cdot \boldsymbol{\nabla} \rho \right) \right) \\ &- \boldsymbol{\zeta}_a \cdot \boldsymbol{\nabla} \left(\boldsymbol{v} \cdot \boldsymbol{\nabla} \rho \right) = - \boldsymbol{v} \cdot \boldsymbol{\nabla} \left(\boldsymbol{\zeta}_a \cdot \boldsymbol{\nabla} \rho \right). \ \end{aligned}$$

using

$$\nabla \cdot \mathbf{v} = \nabla \cdot \zeta_a = 0, \quad \nabla A \cdot (\nabla \wedge B) = -\nabla \cdot (\nabla A \wedge B)$$
$$A \wedge (B \wedge C) = B (A \cdot C) - C (A \cdot B)$$

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Symmetry with respect to Lagrangian relabeling Volume -preserving parcel relabeling in the action principle :

$$oldsymbol{a}
ightarrow oldsymbol{a}' = oldsymbol{a} + \delta oldsymbol{a}(oldsymbol{a},t), \ rac{\partial(oldsymbol{a}',b',c')}{\partial(oldsymbol{a},b,c)} = 1 \Rightarrow$$

$$\boldsymbol{\nabla}_{\boldsymbol{a}} \cdot \delta \boldsymbol{a} = 0 \Rightarrow \delta \boldsymbol{a} = \boldsymbol{\nabla}_{\boldsymbol{a}} \wedge \delta \boldsymbol{\alpha}(\boldsymbol{a}, t). \tag{64}$$

Chain differentiation :

$$\dot{\boldsymbol{X}}\Big|_{\boldsymbol{a}'} = \dot{\boldsymbol{X}}\Big|_{\boldsymbol{a}} + (\boldsymbol{\nabla}_{\boldsymbol{a}} \otimes \boldsymbol{X}) \cdot \dot{\boldsymbol{a}}\Big|_{\boldsymbol{a}'} \Rightarrow \delta \dot{\boldsymbol{X}} = -(\boldsymbol{\nabla}_{\boldsymbol{a}} \otimes \boldsymbol{X}) \cdot \dot{\delta} \boldsymbol{a}.$$

Invariance of action (mass-weighted labels) :

$$\delta S = \int dt \int d^3 \mathbf{a} \, \dot{\mathbf{X}} \cdot \delta \dot{\mathbf{X}} = -\int dt \int d^3 \mathbf{a} \, \dot{\delta \mathbf{a}} \cdot \mathbf{\nu} = 0, \ (65)$$

where $\boldsymbol{\nu} = (\boldsymbol{\nabla}_{\boldsymbol{a}}\otimes \boldsymbol{X})\cdot\dot{\boldsymbol{X}} \equiv (\boldsymbol{\nabla}_{\boldsymbol{a}}\otimes \boldsymbol{X})\cdot\boldsymbol{\nu}$, with duality :

$$\boldsymbol{\nu} = (\boldsymbol{\nabla}_{\boldsymbol{a}} \otimes \boldsymbol{X}) \cdot \boldsymbol{v} \Leftrightarrow \boldsymbol{v} = (\boldsymbol{\nabla}_{\boldsymbol{X}} \otimes \boldsymbol{a}) \cdot \boldsymbol{\nu}.$$
 (66)

Integration by parts in space and time \rightarrow

$$\int dt \int d^{3}\mathbf{a} \frac{d}{dt} \left(\nabla_{\mathbf{a}} \wedge \boldsymbol{\nu} \right) \cdot \delta \boldsymbol{\alpha} = 0 \Rightarrow \frac{d}{dt} \left(\nabla_{\mathbf{a}} \wedge \boldsymbol{\nu} \right) = 0.$$
(67)

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Lagrangian meaning of PV conservation

PV in terms of ν and Lagrangian variables, using (66) :

$$PV := (\boldsymbol{\nabla}_{\boldsymbol{X}} \wedge \boldsymbol{v}) \cdot \boldsymbol{\nabla} \rho = (\boldsymbol{\nabla}_{\boldsymbol{X}} \wedge [\boldsymbol{\nu} \cdot (\boldsymbol{\nabla}_{\boldsymbol{X}} \otimes \boldsymbol{a})]) \cdot \boldsymbol{\nabla} \rho$$

Right-hand side is equivalent to $\sum_{i} \frac{\partial(\nu_{i}, a_{i}, \rho)}{\partial(X, Y, Z)} \equiv \frac{\partial(\nu, a, \rho)}{\partial(X, Y, Z)}$, where i = 1, 2, 3 indicates the x, y, z- components of a vector. Using $\frac{\partial(a, b, c)}{\partial(X, Y, Z)} = \rho_{i}(\mathbf{a})$:

$$PV = \rho_i(\boldsymbol{a}) \frac{\partial(\boldsymbol{\nu}, \boldsymbol{a}, \rho)}{(\boldsymbol{a}, \boldsymbol{b}, \boldsymbol{c})} = \rho_i(\boldsymbol{a}) \left[(\boldsymbol{\nabla}_{\boldsymbol{a}} \wedge \boldsymbol{\nu}) \cdot \boldsymbol{\nabla}_{\boldsymbol{a}} \rho \right],$$

and from (67) and $\dot{
ho}=$ 0, PV is Lagrangian invariant :

$$rac{d}{dt}\left[(oldsymbol{
abla}\wedgeoldsymbol{
abla})\cdot
ho
ight]=0$$

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Stationary solutions, jets

- Trivial stationary solution : a state of rest in hydrostatic equilibrium.
- Non-trivial stationary solutions :
 - β- plane : zonal flows u = U(y, z) in geostrophic and hydrostatic equilibrium

$$-f(y)U(y,z) = -\partial_y \Phi(y,z), \ v = w = \sigma = 0, \quad (69)$$

$$g\frac{\rho_s(y,z)}{\rho_0} = -\partial_z \Phi(y,z) \tag{70}$$

► *f*- plane : arbitrary oriented (zonal or meridional) flows. Jets : localized $U(y, z) \rightarrow$ nonzero PV anomaly $(-\partial_y U)(\partial_z \rho_s) \neq 0.$ Mathematics of the atmosphere and oceans 1

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Polar coordinates

$$(x, y, z) \rightarrow (r, \theta, z)$$

Horizontal velocity :

$$oldsymbol{v}_h = (uoldsymbol{\hat{r}} + voldsymbol{\hat{ heta}})$$

Horizontal divergence :

$$\boldsymbol{\nabla}_h \cdot \boldsymbol{v}_h = \frac{1}{r} \left(\partial_r (ru) + \partial_\theta v \right)$$

Lagrangian derivative :

$$\frac{d}{dt} = \partial_t + u\partial_r + \frac{v}{r}\partial_\theta + w\partial_z$$

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Stationary solutions : vortices Primitive equations for horizontal motion

$$\frac{d\boldsymbol{v}_h}{dt} + \left(f + \frac{v}{r}\right)\hat{z} \wedge \boldsymbol{v}_h + \boldsymbol{\nabla}_h \phi = 0, \quad (71)$$

Vortex solutions

Stationary axisymmetric solution on the *f*- plane : cyclo-geostrophic + hydrostatic equilibria :

 $u = w = \sigma = 0, v = V(r, z)$

$$\left(f + \frac{V(r,z)}{r}\right)V(r,z) = \partial_r \Phi(r,z)$$
 (73)

$$g\frac{\rho_{\mathfrak{s}}(r,z)}{\rho_{0}} = -\partial_{z}\Phi(r,z) \tag{74}$$

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Small perturbations over the state of rest on the *f*- plane

Linearized equations :

Perturbations about the state of rest : $\mathbf{v} = 0$, with linear stratification on the *f*- plane, f = const. Linearized equations :

$$u_{t} - fv + \phi_{x} = 0,$$

$$v_{t} + fu + \phi_{y} = 0,$$
 (75)

$$\phi_{z} + \frac{g}{\rho_{0}}\sigma = 0, \quad \sigma_{t} + w\rho'_{s} = 0,$$

$$u_{x} + v_{y} + w_{z} = 0,$$
 (76)

where u, v, w - three components of velocity perturbation, ϕ - geopotential perturbation, σ - perturbation of density, ρ_s - background density $\rho'_s = const$. Notation : $(...)_x := \partial_x(...)$ etc. Mathematics of the atmosphere and oceans 1

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Dispersion relation for Fourier-mode solutions Elimination of σ and w:

$$u_{t} - fv + \phi_{x} = 0,$$

$$v_{t} + fu + \phi_{y} = 0,$$

$$u_{x} + v_{y} - N^{-2}\phi_{zzt} = 0,$$
(77)

a system of PDEs with constant coefficients. $N^2 = -\frac{g\rho'_s}{\rho_0}$ - Brunt - Väisälä frequency.

 $\mathsf{Fourier}\text{-}\mathsf{transform}\leftrightarrow \mathsf{harmonic}\ \mathsf{waves}:$

$$(u, v, \phi) = (u_0, v_0, \phi_0)e^{i(\omega t - \mathbf{k} \cdot \mathbf{x})} + \text{c.c.} \Rightarrow \qquad (78)$$

System of homogeneous algebraic equations for (u_0, v_0, ϕ_0) , with solvability condition giving dispersion equation

$$\omega \left(\omega^2 - \left(N^2 \frac{k_x^2 + k_y^2}{k_z^2} + f^2 \right) \right) = 0.$$
 (79)

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Physical meaning of solutions

Three roots of this equation provide :

 Stationary solutions ω = 0; correspond to linearized PV conservation equation :

$$\partial_t q_L = 0, \quad q_L = -(v_x - u_y) \frac{\rho_0}{g} N^2 + f \sigma_z, \qquad (80)$$

and hence to vortex motions.

Propagative waves with dispersion relation :

$$\omega = \pm \sqrt{N^2 \frac{k_x^2 + k_y^2}{k_z^2} + f^2}$$
 (81)

and two senses of propagation, internal inertia-gravity waves (IGW). They are supra-inertial : $|\omega| \ge f$.

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Waves on the background of a jet

Linearized equations :

Perturbations about a stationary jet : u = U(y, z), v = w = 0, with stratification $\rho_s(z)$. Linearized equations :

$$u_{t} + Uu_{x} + vU_{y} + wU_{z} - fv + \phi_{x} = 0,$$

$$v_{t} + Uv_{x} + fu + \phi_{y} = 0,$$
 (82)

$$\phi_{z} + \frac{g}{\rho_{0}}\sigma = 0, \quad \sigma_{t} + U\sigma_{x} + w\rho'_{s} = 0,$$

$$u_{x} + v_{y} + w_{z} = 0,$$
 (83)

Fourier-transformation \rightarrow waves propagating along the jet :

$$(u, v, \phi) = (u_0, v_0, \phi_0)(y, z)e^{i(\omega t - k x)} + c.c.,$$
 (84)

In general, complex eigen-frequencies $\omega \Rightarrow$ growing eigen-modes \rightarrow instability of the jet.

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Full linearized non-hydrostatic equations

$$g\frac{\sigma}{\rho_0} = -\phi_z \to w_t + g\frac{\sigma}{\rho_0} = -\phi_z.$$
 (85)

Elimination of buoyancy $b = g \frac{\sigma}{\rho_0}$ and vertical velocity w :

$$b = -\phi_z - w_t, \quad -\left(\partial_{tt} + N^2\right)\left(u_x + v_y\right) + \phi_{zzt} = 0 \Rightarrow (86)$$

$$u_{t} - fv = -\phi_{x}, \qquad (87)$$

$$v_{t} + fu = -\phi_{y}, \qquad (88)$$

$$(\partial_{tt} + N^{2}) (u_{x} + v_{y}) - \phi_{zzt} = 0, \qquad (89)$$

Dispersion relation :

$$\omega \left[\omega^2 - \left(N^2 \frac{k^2 + l^2}{k^2 + l^2 + m^2} + f^2 \frac{m^2}{k^2 + l^2 + m^2} \right) \right] = 0 \quad (90)$$

Typically in the atmosphere and ocean

$$N^2 > f^2 \Rightarrow f^2 \le \omega^2 \le N^2$$
 (91)

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Introduction

What is GFD ? Vortices, jets, and waves

Hydrodynamics, a refresher

Equations of the perfect fluid Geometric view of the perfect fluid Dissipation

Primitive equations (PE)

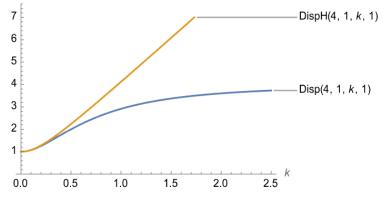
Rotation, sphericity, tangent plane approximation PE : Ocean PE : Atmosphere

Vortices, jets, and waves in PE

Conservation of potential vorticity (PV) Stationary jets and vortices Linear waves

What we lose imposing hydrostatics

Hydrostatic vs non-hydrostatic dispersion curves



Frequency as a function of the modulus of horizontal wavenumber k at fixed vertical wavenumber m = 1, and N = 4, f = 1 as given by hydrostatic (DispH) vs non-hydrostatic (Disp) dispersion relations.

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What we lose imposing hydrostatics

- Large-scale atmospheric and oceanic motions : same set of PE, modulo changes of variables
- Crucial property of PE : Lagrangian conservation of PV.
- ► Main dynamical entities in PE : vortices, jets, and waves
- Vortices : slow motions carrying non-zero PV anomaly, have zero frequency in linear approximation,
- Jets : particular case of vortex motions.
- Inertia-gravity waves : fast motions bearing no PV anomaly
- Frequencies of wave and vortices are separated by a spectral gap [0, f], frequencies of the waves bounded from below by f.
- Non-hydrostatic effects, if included, impose an upper boundary to the wave spectrum.

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